A satellite image from Terra MODIS showing a large dust plume originating from the coast of North Africa, specifically near Benghazi, and extending into the Mediterranean Sea. The land is shown in shades of brown and tan, while the sea is dark blue. The dust plume is a lighter, hazy yellowish-brown color. Labels for 'Mediterranean Sea', 'Benghazi', and 'Gulf of Sidra' are visible. A scale bar for 25 km and a north arrow are in the bottom right.

Predicting the Mineral Composition of Dust Aerosols With an Earth System Model (NASA GISS ModelE2)

Jan P. Perlwitz, Carlos Pérez Garcia-Pando, and Ron L. Miller

Department of Applied Physics and Applied Mathematics, Columbia University, and NASA Goddard Institute For Space Studies, New York, USA, Contact: jan.p.perlwitz@nasa.gov

Acknowledgements: The work has been supported by the Department of Energy, NASA, NSF, and Ministry Economy and Competitiveness of Spain

Image: Terra MODIS, September 28, 2010. NASA Image by Jeff Schmaltz, MODIS Rapid Response Team at NASA GSFC 25 km 

Outline

1. Soil (mineral) dust aerosols in Earth's climate system
2. Globally uniform soil dust in ModelE
3. Predicting the mineral composition of dust aerosols in ModelE
4. Outlook to next steps and near future work

Soil Dust Aerosols in the Climate System

- Major aerosol in the atmosphere (global emission 1000-3000 Tg/a; total mass load in atmosphere approximately 20-35 Tg)
- High spatial and temporal variability
- Absorbs and reflects radiation => impact on radiation balance, warming or cooling effect depending on the single scattering albedo of the dust particles
- Carrier of nutrients like iron => Fertilization of phytoplankton
=> Carbon cycle
- Cloud condensation nuclei, ice nuclei
- Atmospheric chemistry (e.g., uptake of SO_2 , H_2SO_4 , HNO_3 , N_2O_5 , NO_3 , NO_2 , O_3 , H_2O_2 , OH , HO_2) => formation of coatings, like sulfates, nitrates, and/or effect on trace gas budgets

Dust Particles

4722

A. Alastuey et al. / Atmospheric Environment 39 (2005) 4715–4728

(Alastuey et al., 2005)

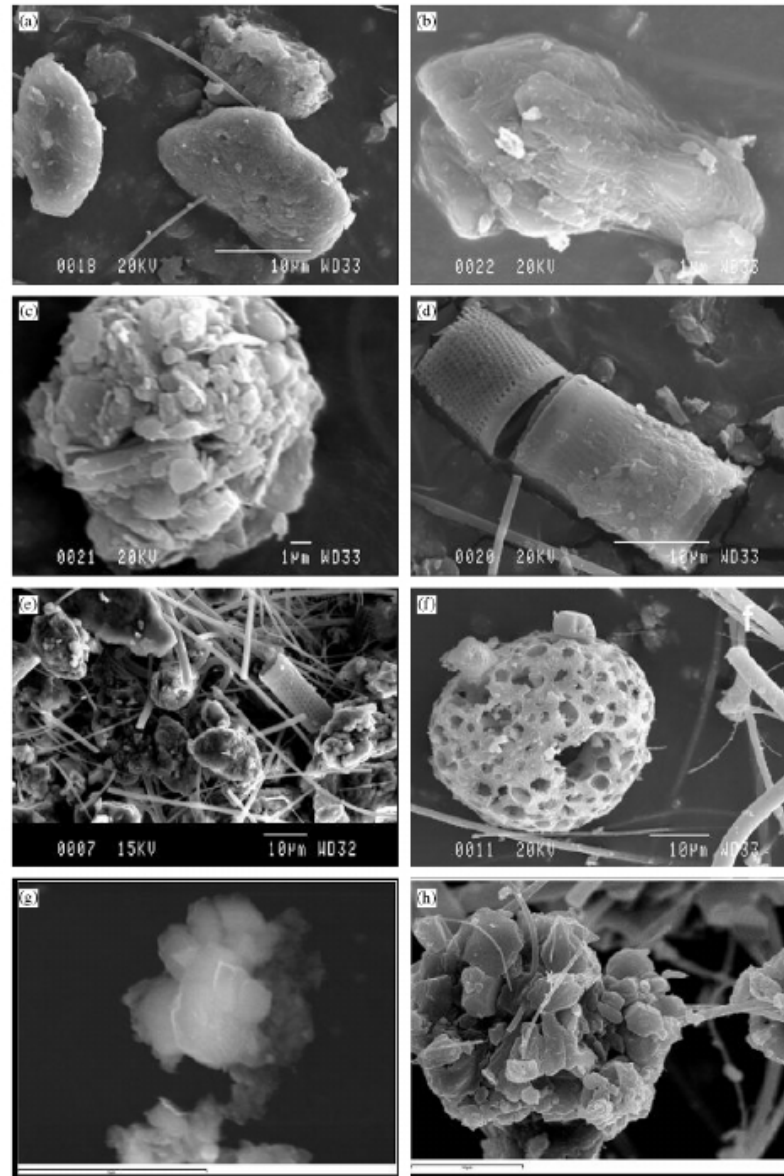


Fig. 5. SEM microphotographs of TSP collected the 29th July 2002 at IZO (a–d) and at SCO (e–h). IZO: (a) rounded quartz; (b) large (>10 μm) plate clay particles; (c) aggregates of micro-crystals of clay minerals; (d) silica skeletons of fresh water diatom from the *Melosira* genus, present in lakes or ponds from Northern Africa. SCO: (e) general aspect showing the presence of mineral dust, marine aerosol and *Melosira* genus diatoms, coated with sulphate; (f) spongy carbonaceous particles with sodium chloride crystals from marine aerosols; (g) K/Ca sulphate coating clay aggregates and (h) <5 μm Ca sulphate particles with crystalline habit.

Global Dust Sources

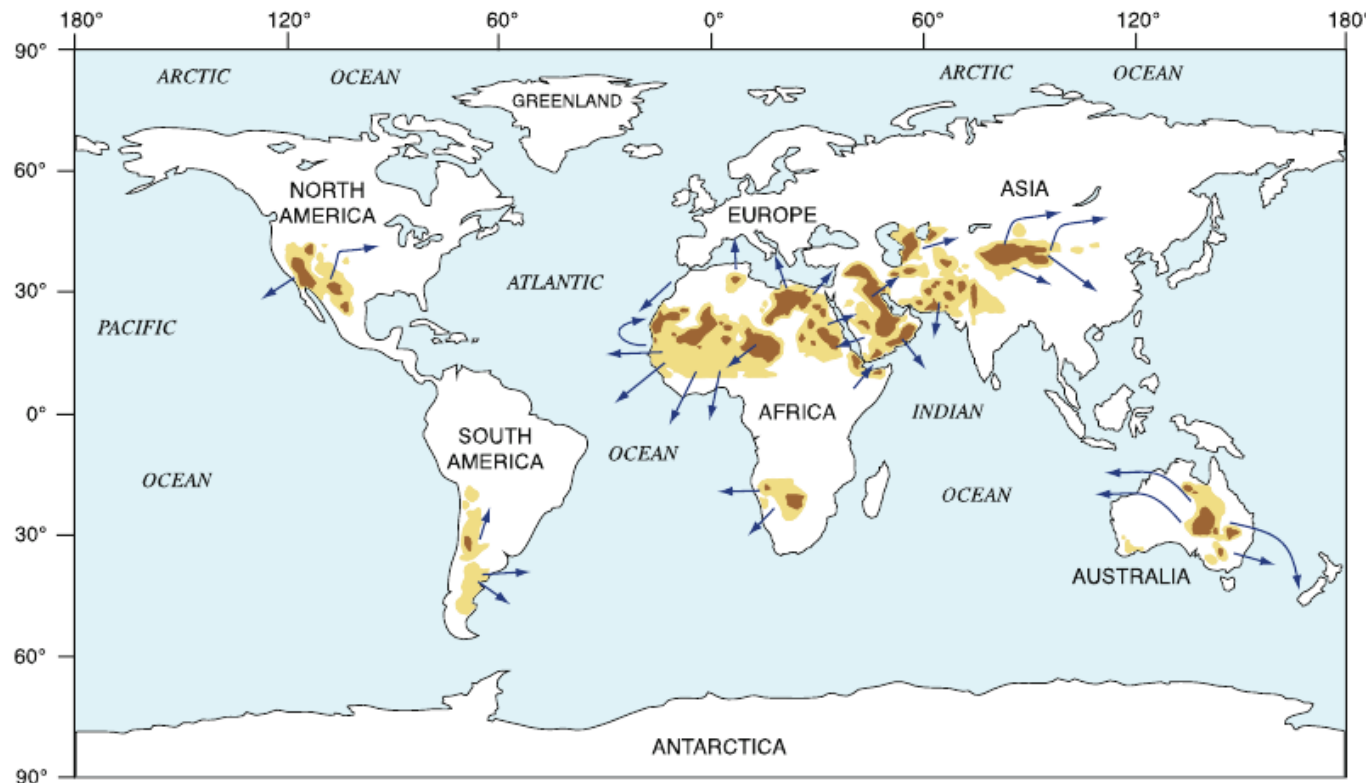


Fig. 3.2 Map of global dust sources, based on multiple years of satellite imagery, derived from frequency of occurrence (number of days) where the TOMS absorbing aerosol index (AAI) is greater than 0.7 (significant amounts of dust or smoke) or 1.0 (abundant dust or smoke). For comparison, nonabsorbing aerosols such as sulfate and sea salt yield negative AAI values; clouds yield values near zero; ultraviolet-absorbing aerosols such as dust and smoke yield positive values (Prospero et al. 2002). *Dark brown* is 21–31 days; *yellow* is 7–21 days (Redrawn from Fig. 4 of Prospero et al. 2002). *Blue arrows* indicate typical dust transport pathways, based on interpretation of MODIS imagery from Terra and Aqua satellites by the authors

(Muhs et al. 2014)

Global Dust Emission:
1000-3000 Tg/a
(Cakmur et al. 2006)

Northern Africa:
515 Tg/a
(Miller et al. 2004)
to 1087 Tg/a
(Tanaka and Chiba, 2006)

East Asia:
54 Tg/a (Luo et al., 2003)
to 460 Tg/year
(Laurent et al. 2006)

Arabian Peninsula:
43 Tg/a
(Miller et al., 2004)
to 496 Tg/a
(Ginoux et al., 2004)

Australia:
37 Tg/a
(Zender et al., 2003)
to 148 Tg/a
(Miller et al., 2004)

Mie Scattering and Absorption by Mineral Dust

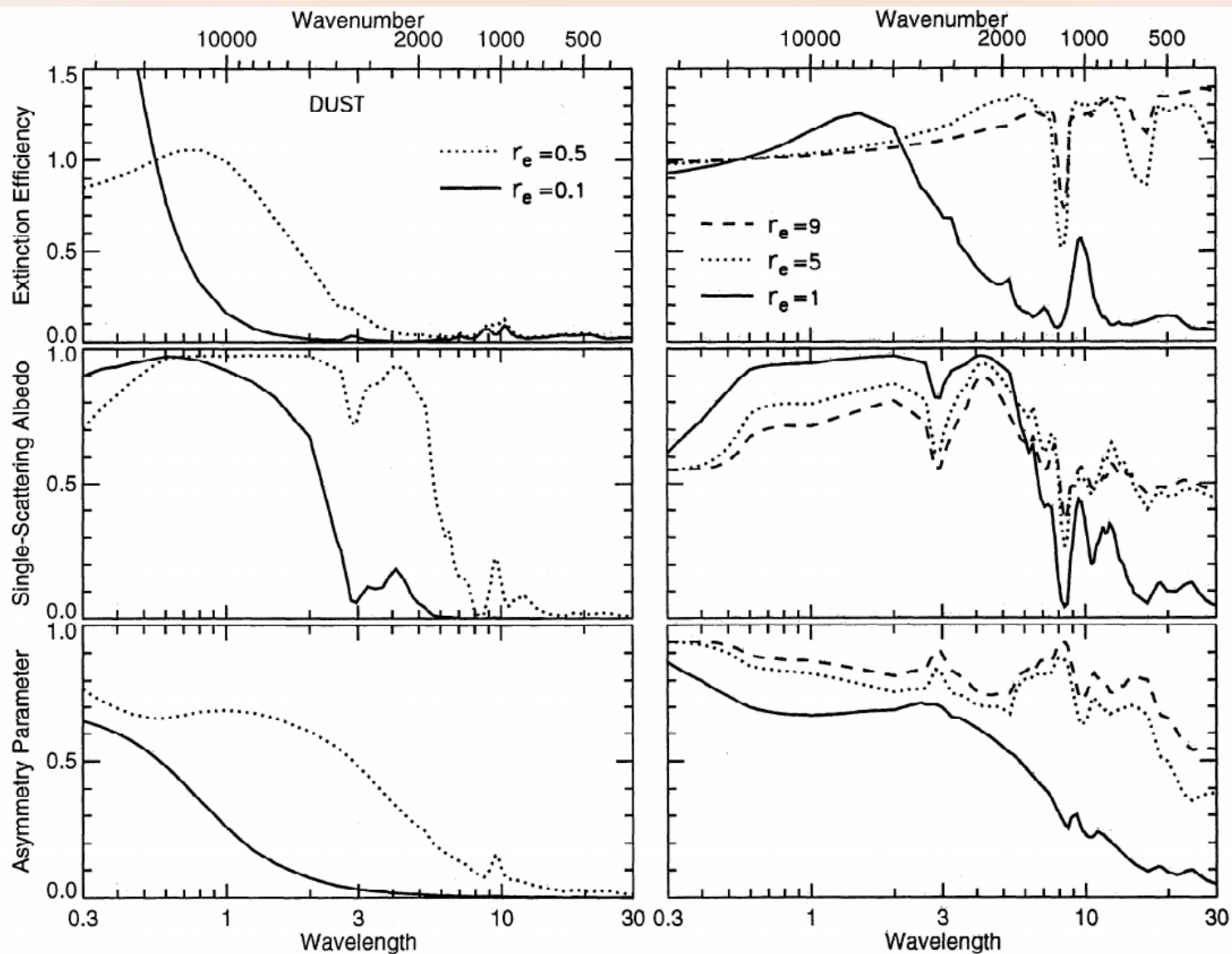


Figure 1. Mie scattering calculated radiative parameters for desert dust aerosol for eight particle sizes (effective radii, r_e , given in micrometers) in dependence on wavelength (in micrometers).

(Tegen and Lacis, JGR 1996)

Uncertainty in Radiative Properties of Mineral Dust

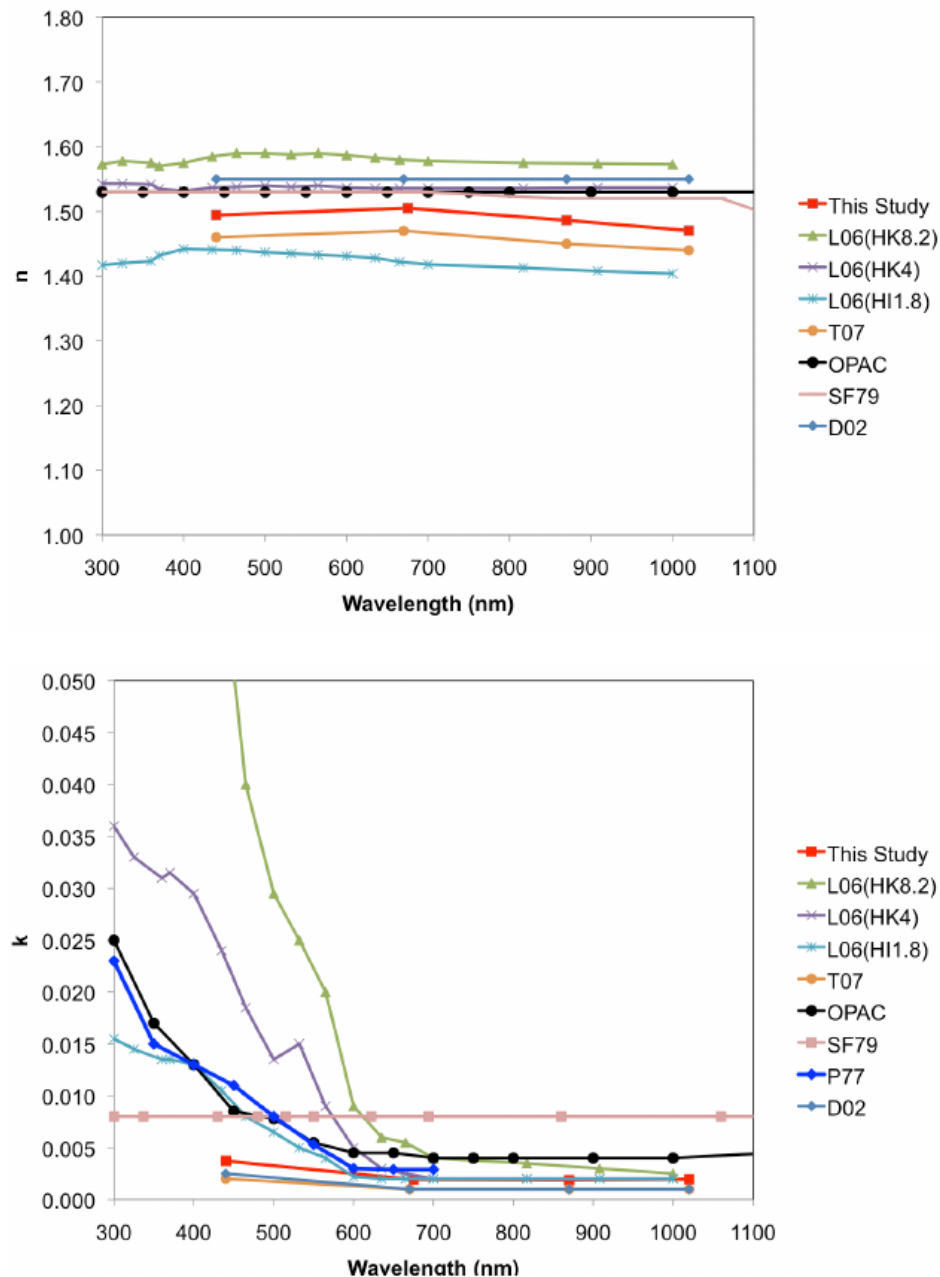
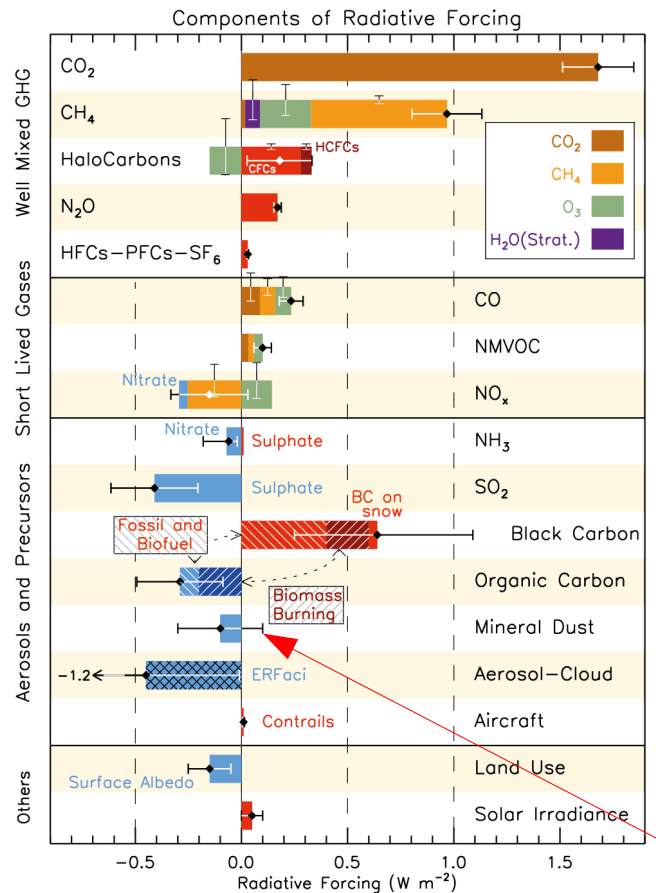


Fig. 5. Real and imaginary refractive index of the dust determined from AERONET (DU) from the current study. T07 is from Todd et al. (2007), OPAC is from Hess et al. (1998), SF79 is from Shettle and Fenn (1979), L06 is from Lafon et al. (2006), and D02 is the Bahrain site (most absorbing site) in Dubovik et al. (2002). Different volume fractions (%) of iron oxide in the hematite-kaolinite (HK) and hematite-illite (HI) aggregates are shown in the parenthesis of L06.

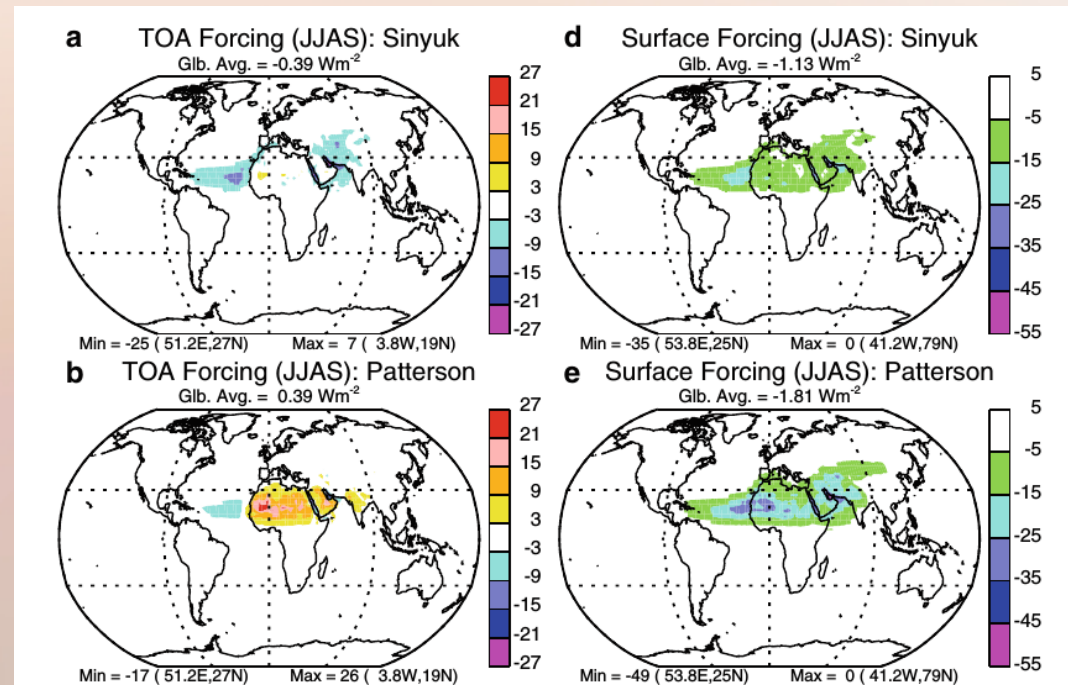
(Kim et al., ACP 2011)

Radiative Forcing by Soil Dust



(IPCC, 2013)

Mineral Dust



(Miller et al., 2014)

- Radiative forcing: Change in radiation flux before climate response
- Globally averaged top of atmosphere (TOA) radiative forcing by dust is relatively small, -0.4 to 0.4 W/m^2 (pre-industrial to present day, -0.3 to 0.1 W/m^2 , for comparison TOA forcing by CO_2 : ca. 1.9 W/m^2), globally averaged dust surface forcing always negative, larger in magnitude, $< -1 \text{ W/m}^2$
- However, forcing by dust regionally much larger

Impact on Human Health

All EHP content is accessible to individuals with disabilities. A fully accessible (Section 508-compliant) HTML version of this article is available at <http://dx.doi.org/10.1289/ehp.1306640>.

Research

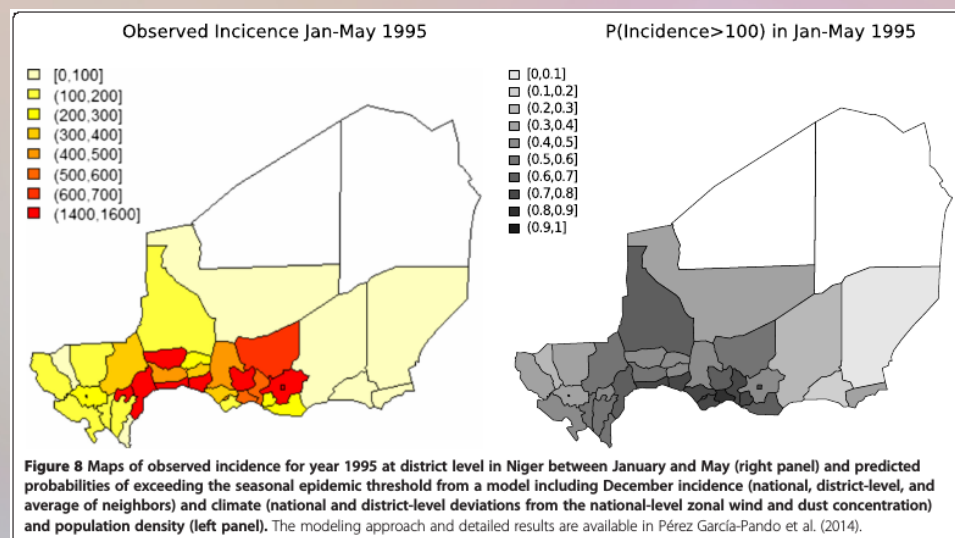
Soil Dust Aerosols and Wind as Predictors of Seasonal Meningitis Incidence in Niger

Carlos Pérez García-Pando,^{1,2*} Michelle C. Stanton,^{3,4,*} Peter J. Diggle,^{3,5} Sylwia Trzaska,⁶ Ron L. Miller,^{1,2} Jan P. Perlwitz,^{1,2} José M. Baldasano,⁷ Emilio Cuevas,⁸ Pietro Ceccato,⁶ Pascal Yaka,⁹ and Madeleine C. Thomson^{6,10}

¹NASA Goddard Institute for Space Studies, New York, New York, USA; ²Department of Applied Physics and Applied Math, Columbia University, New York, New York, USA; ³Lancaster Medical School, Lancaster University, Lancaster, United Kingdom; ⁴Liverpool School of Tropical Medicine, Liverpool, United Kingdom; ⁵Department of Epidemiology and Population Health, University of Liverpool, Liverpool, United Kingdom; ⁶International Research Institute for Climate and Society, Palisades, New York, USA; ⁷Barcelona Supercomputing Center–Centro Nacional de Supercomputación, Barcelona, Spain; ⁸Izaña Atmospheric Research Center, Agencia Estatal de Meteorología, Tenerife, Spain; ⁹Office of Civil Aviation and Meteorology General Direction, Ouagadougou, Burkina Faso; ¹⁰Department of Environmental Health Sciences, Mailman School of Public Health, Columbia University, New York, New York, USA

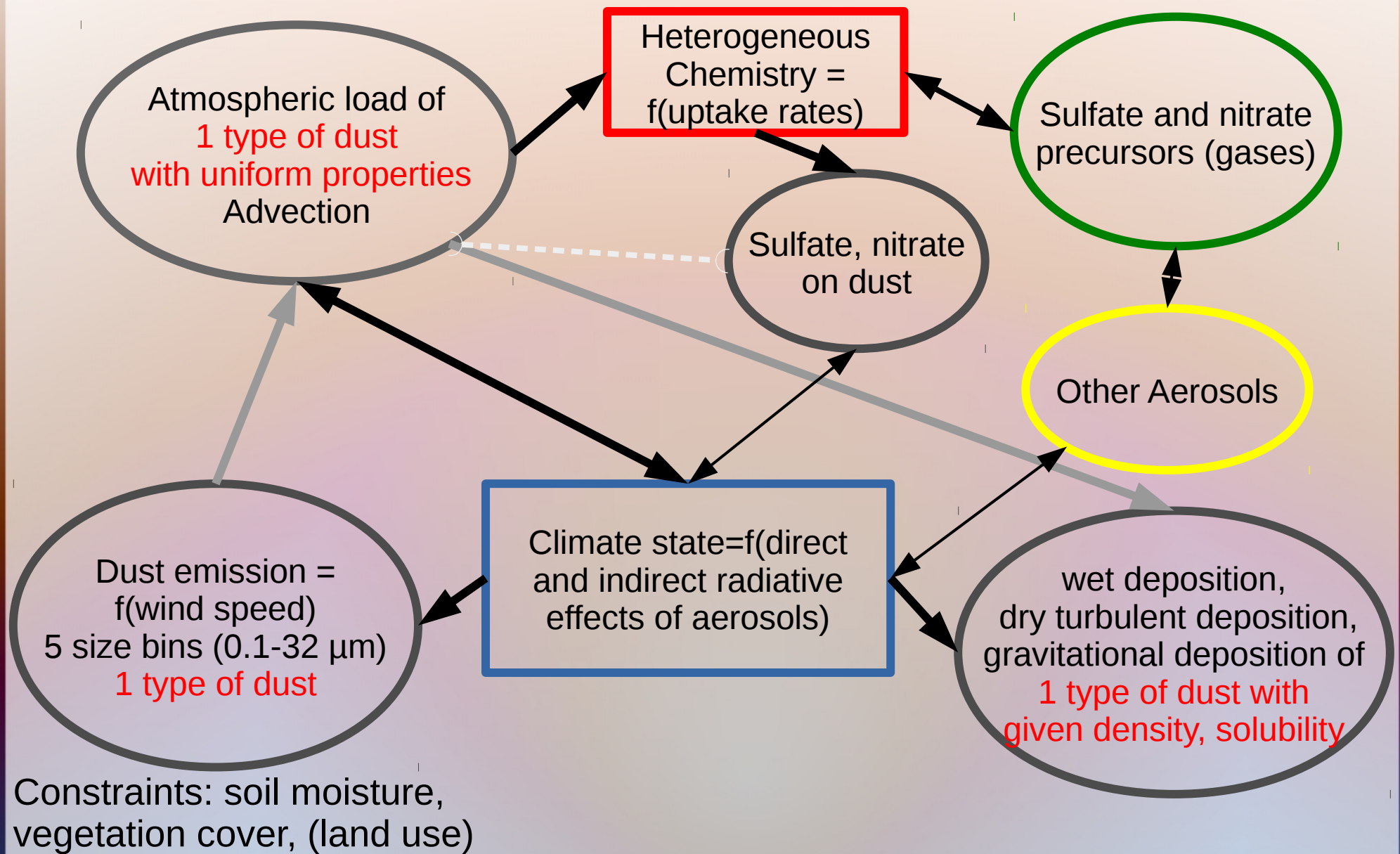
*These authors contributed equally.

(Pérez García-Pando et al., EHP 2014)



(Pérez García-Pando et al., EP 2014)

Uniform Soil Dust in NASA-GISS ModelE2



Probability Scheme for Dust Emission

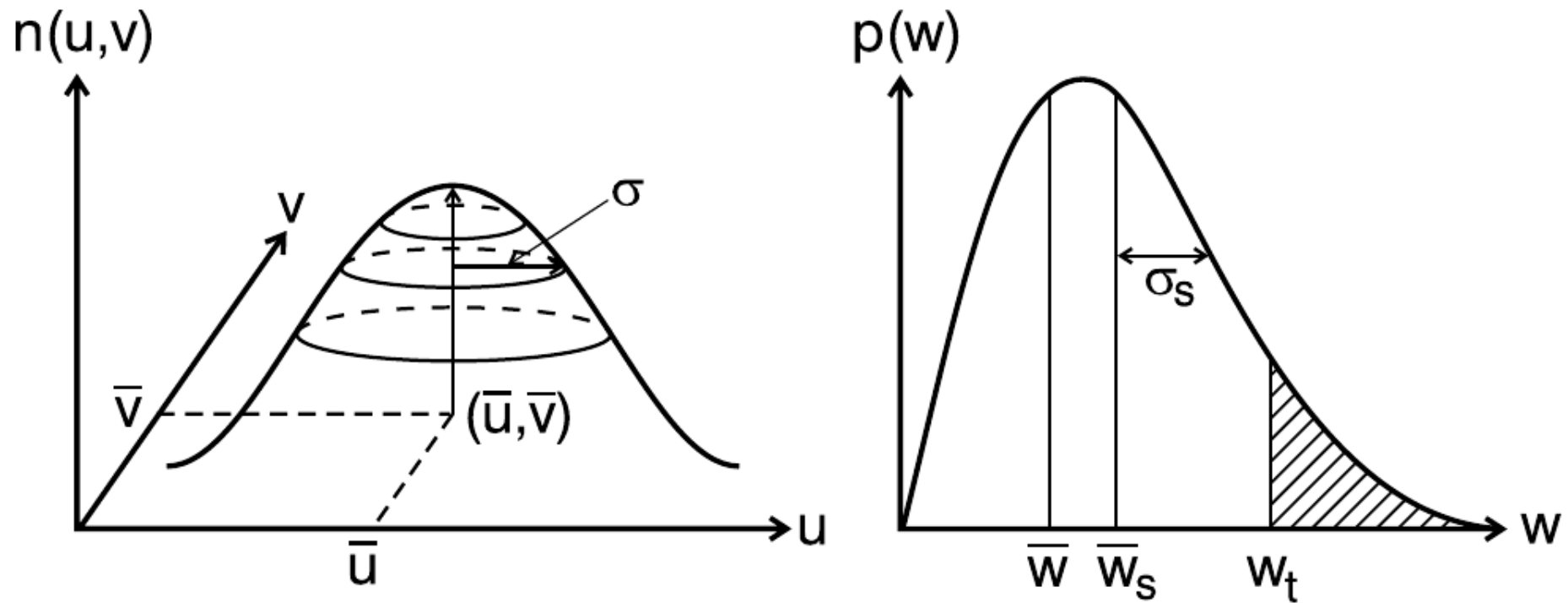


Figure 5. Probability functions representing the distribution of (left) wind velocity and (right) wind speed within the model grid box. Note that the mean wind speed $\bar{w}_s \equiv \int_0^\infty wp(w)dw$ is distinct from the speed of the mean wind $\bar{w} \equiv \sqrt{\bar{u}^2 + \bar{v}^2}$ that is explicitly calculated by the atmospheric general circulation model (AGCM). Similarly, the standard deviation of the wind speed, $\sigma_s \equiv \int_0^\infty (w - \bar{w}_s)^2 p(w)dw$, differs from σ , the standard deviation of each component of the wind velocity.

Globally Uniform Dust in ModelE (used up to CMIP5)

Optical Depth

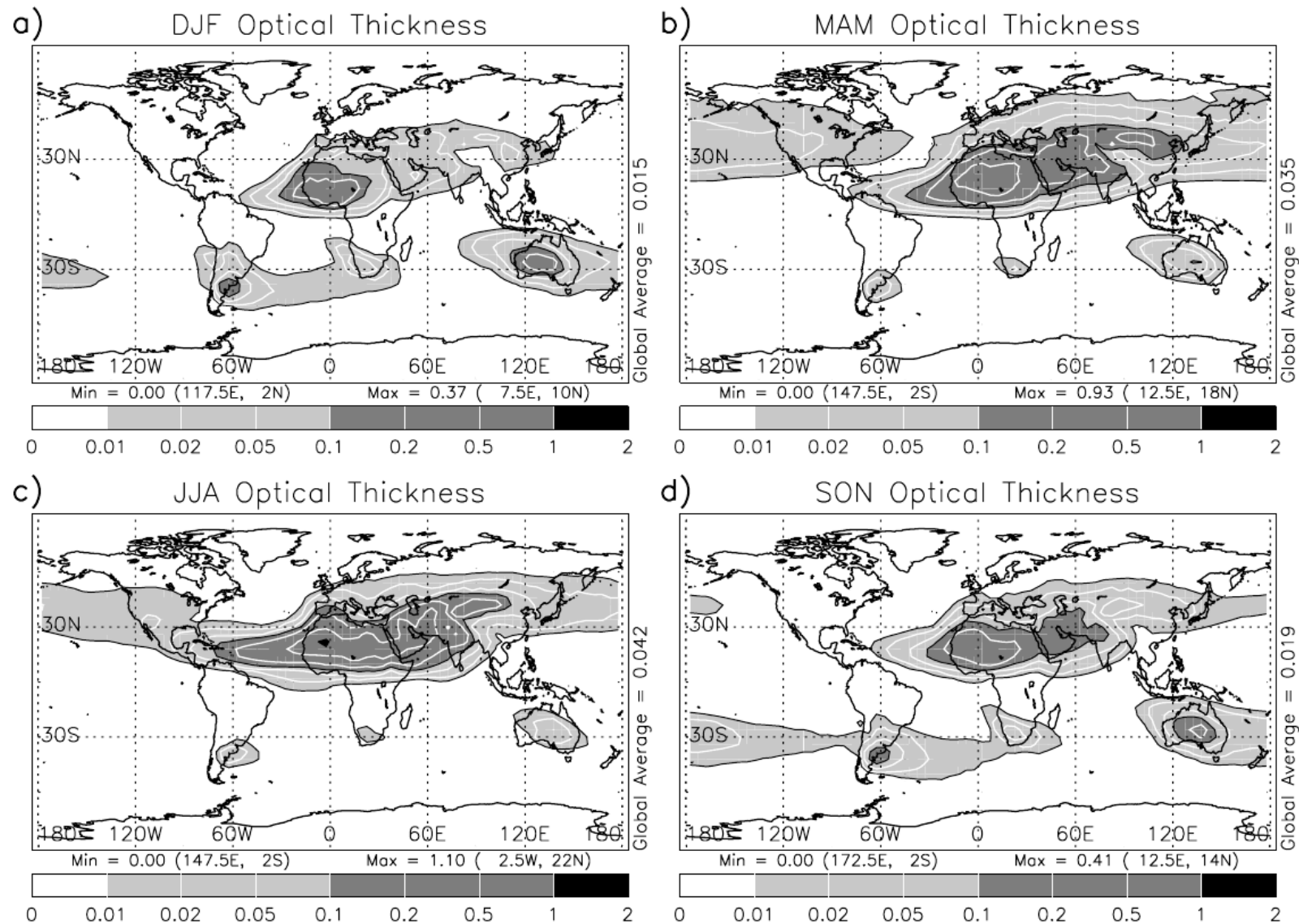


Figure 4. Clear-sky dust optical thickness for (a) DJF, (b) MAM, (c) JJA, and (d) SON.

(Miller et al., JGR 2006)

Globally Uniform Dust in ModelE (used up to CMIP5)

Single Scattering Albedo

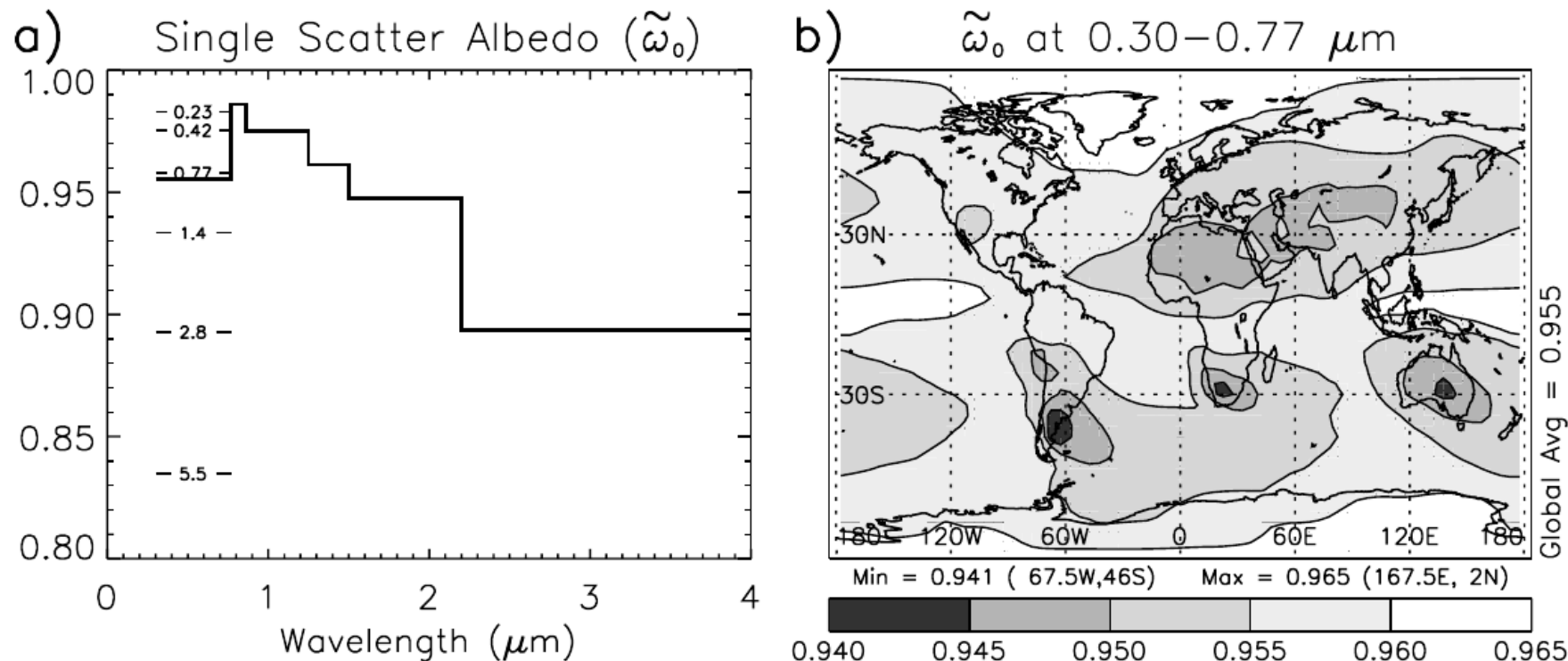


Figure 3. (a) Annual and global average of the bulk single scatter albedo ϖ_0 , for the six spectral bands used in the calculation of solar radiative fluxes. The single scatter albedo is computed for each month as the ratio of the column scattering extinction and the column total extinction by dust. Also shown for the shortest solar band (0.30–0.77 μm) is the single scatter albedo for each individual particle size category, labeled by its effective radius. (b) Geographic distribution of the annual average bulk single scatter albedo at 0.3–0.77 μm .

(Miller et al., JGR 2008)

Why is the Mineral Composition of Soil Dust Important?

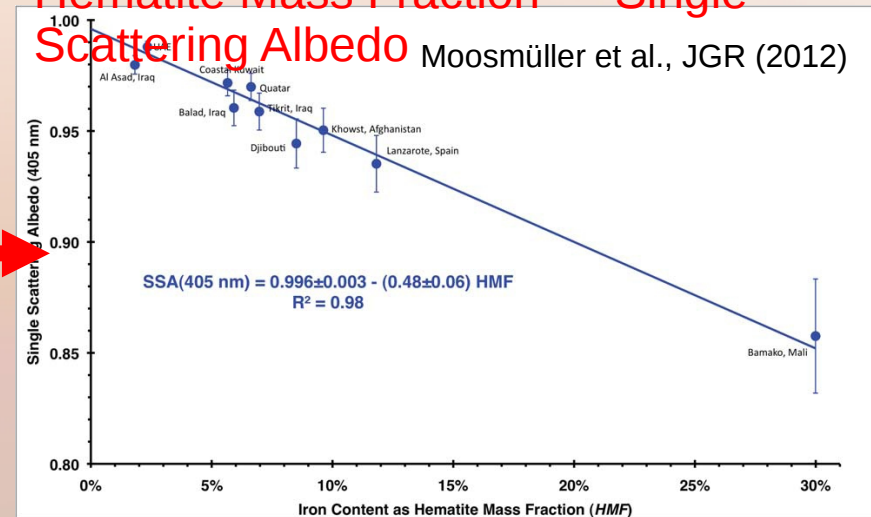
- Traditionally, global dust models have used globally uniform dust properties

Limitations for following reasons:

- Single scattering albedo (SSA) of dust particles depends on mineral composition, particularly on the mass fraction of hematite or goethite
- In turn, aerosol forcing and the response of clouds and atmospheric circulation to the forcing depend on the SSA
- Heterogenous chemistry (e.g., uptake rates) of dust particles depends on mineralogical and chemical composition
- Hygroscopicity of dust particles, the ability to act as cloud condensation nuclei, depends on the mineralogical composition
- Fertilization of phytoplankton in oceans is linked to availability of soluble iron, i.e., to the mineral types of dust

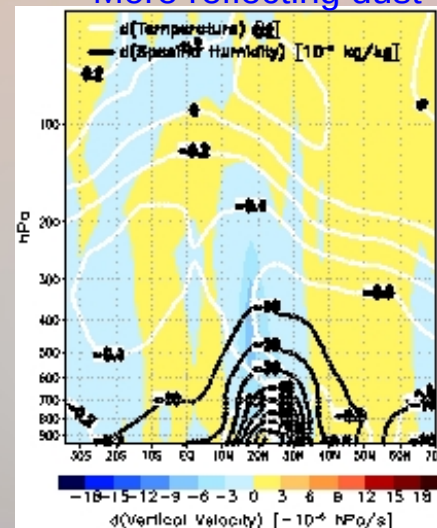
Hematite Mass Fraction → Single Scattering Albedo

Moosmüller et al., JGR (2012)

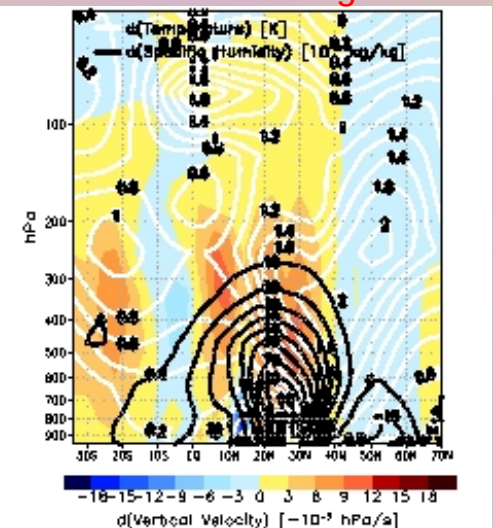


Perlwitz and Miller, JGR (2010),
 Zonal Average 20° – 85° E in JJA:

More reflecting dust

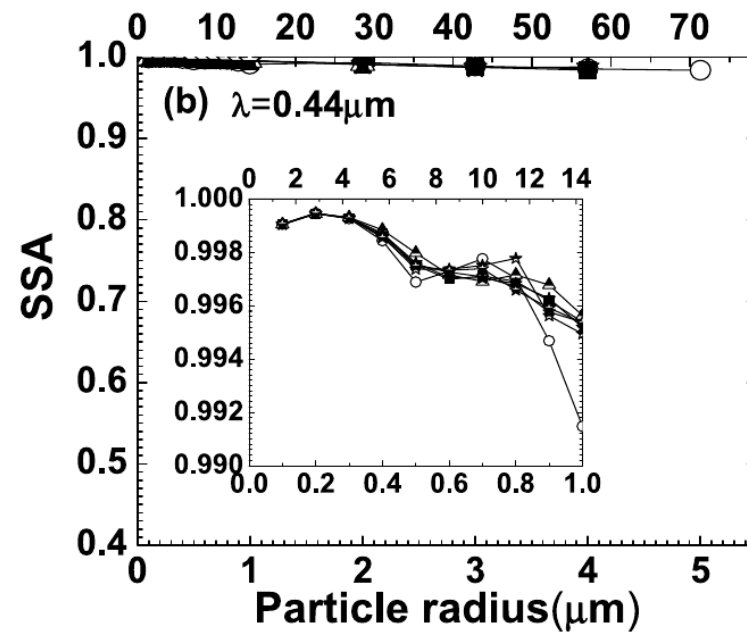
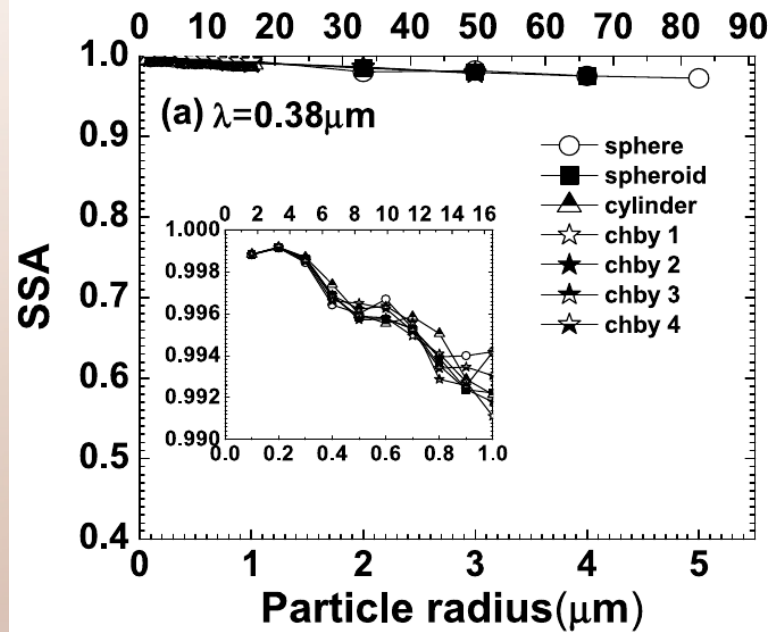


More absorbing dust

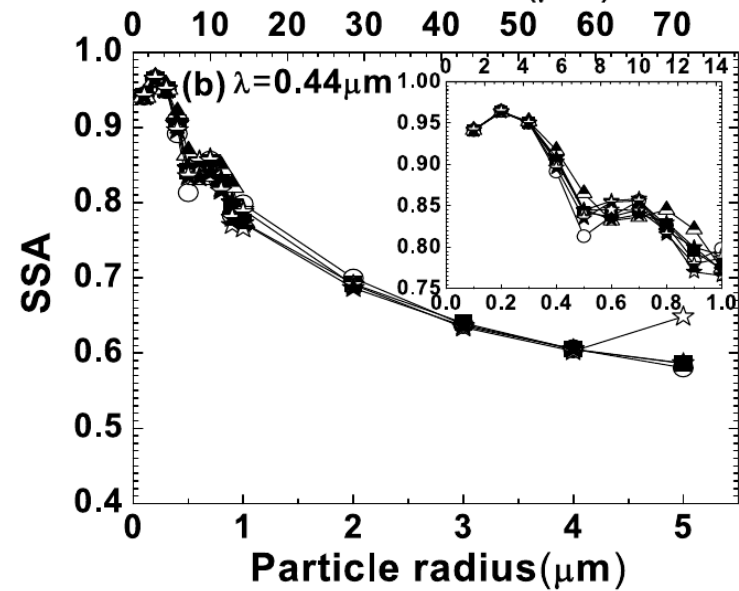
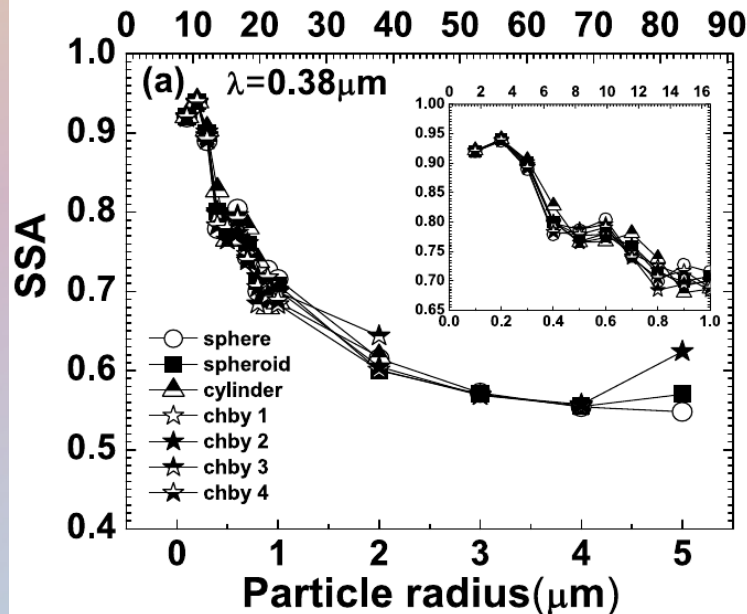


Shaded: Blue: Less Upward Red: More Upward

Hematite Content in Dust and Absorptivity



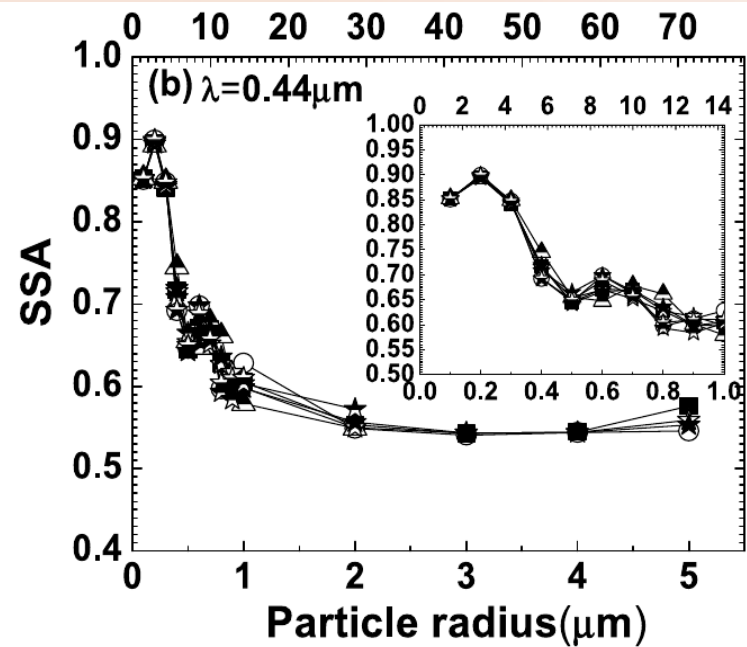
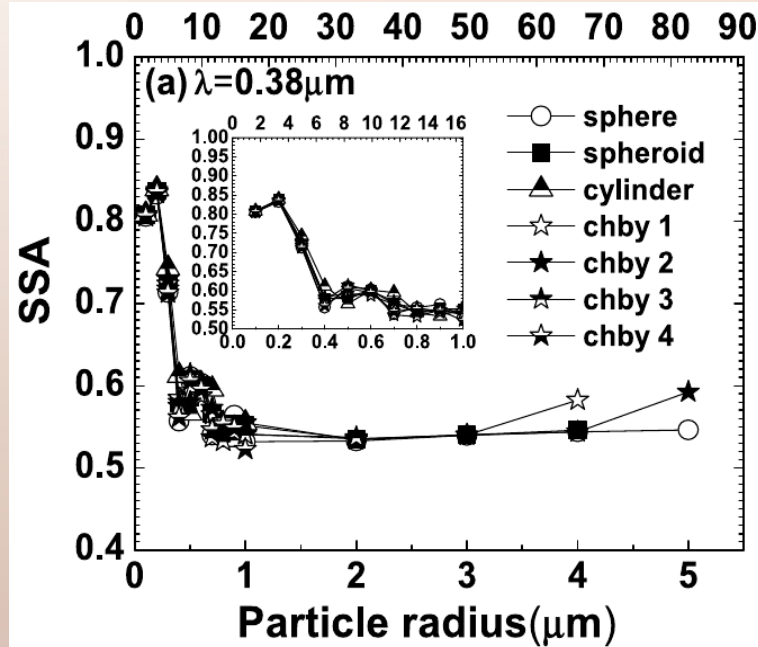
0%



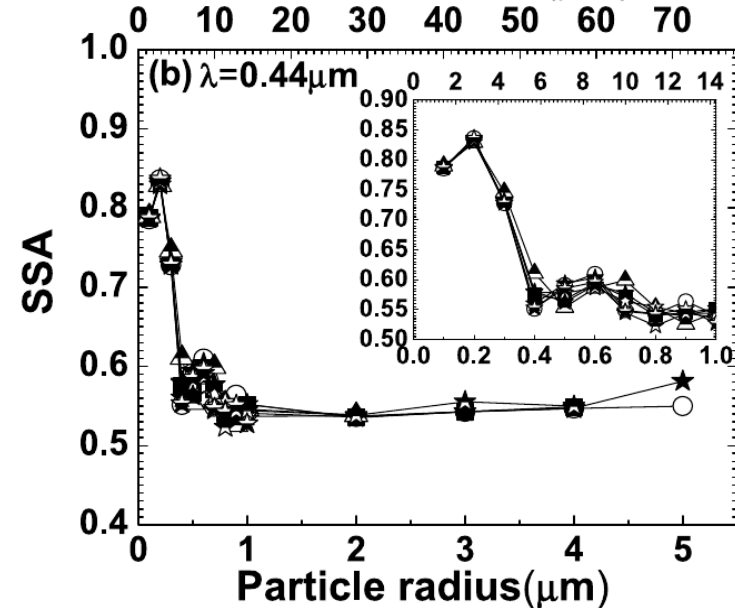
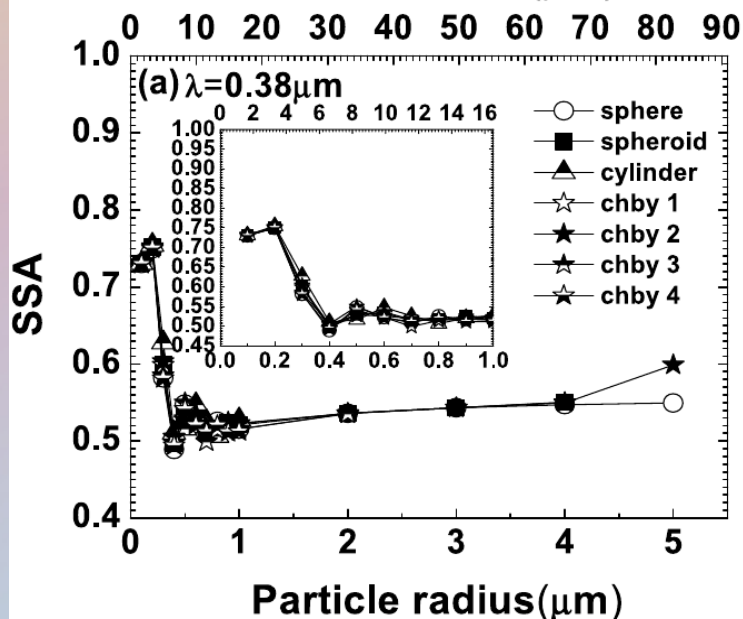
2%

(Mishra and Tripathi, JGR 2008)

Hematite Content in Dust and Absorptivity



6%



10%

(Mishra and Tripathi, JGR 2008)

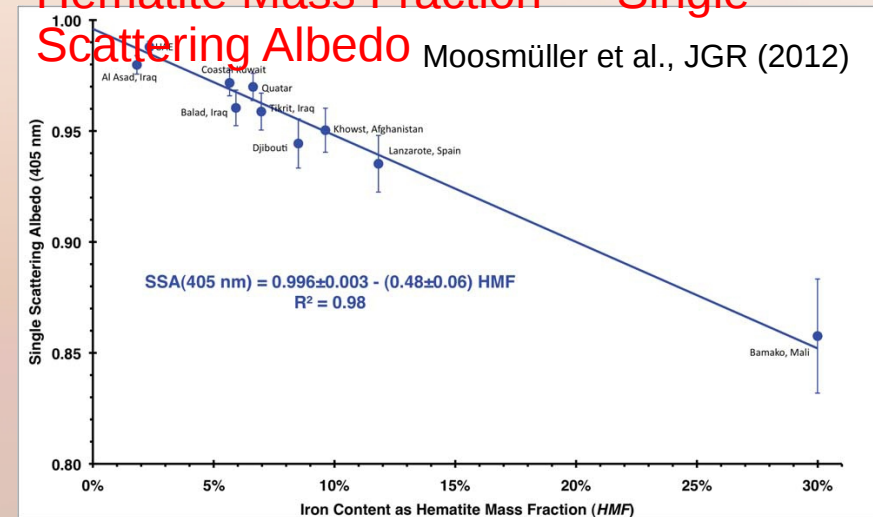
Why is the Mineral Composition of Soil Dust Important?

- Traditionally, global dust models have used globally uniform dust properties

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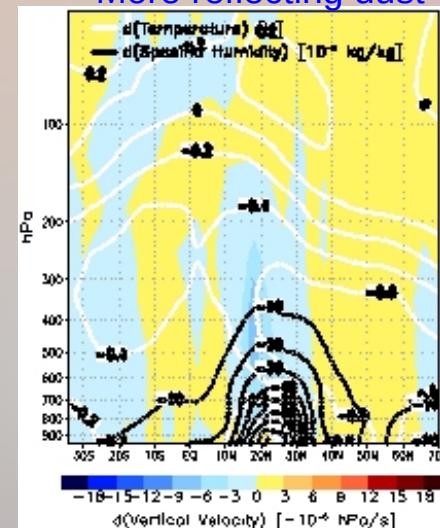
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Hematite Mass Fraction → Single Scattering Albedo

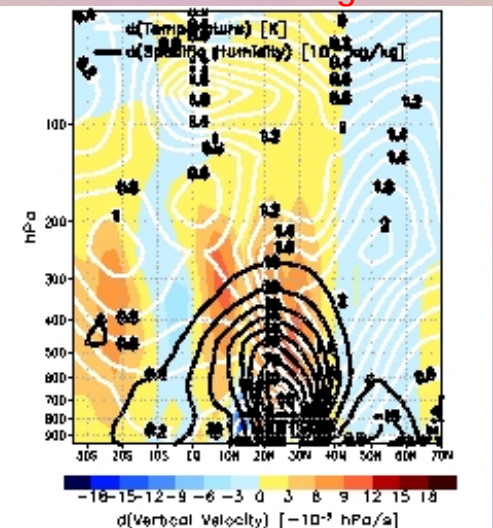


Perlwitz and Miller, JGR (2010),
 Zonal Average 20° – 85° E in JJA:

More reflecting dust



More absorbing dust



Shaded: Blue: Less Upward Red: More Upward

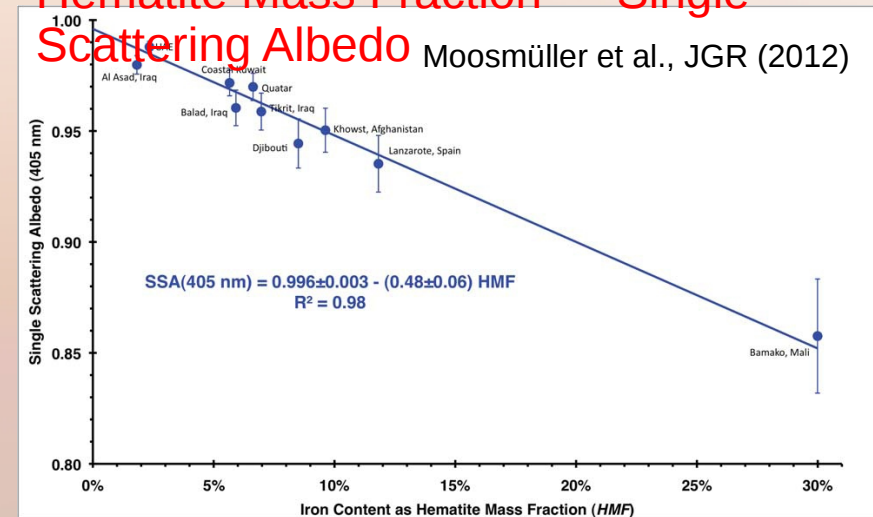
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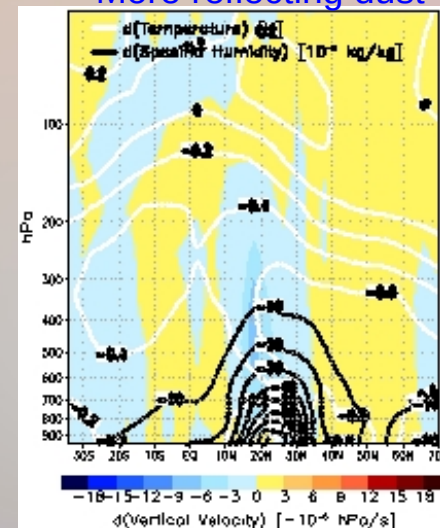
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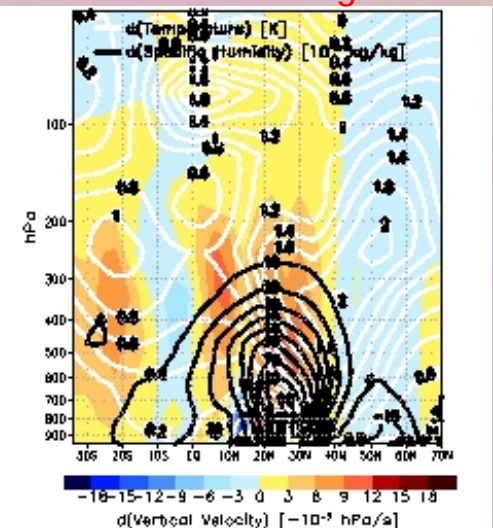


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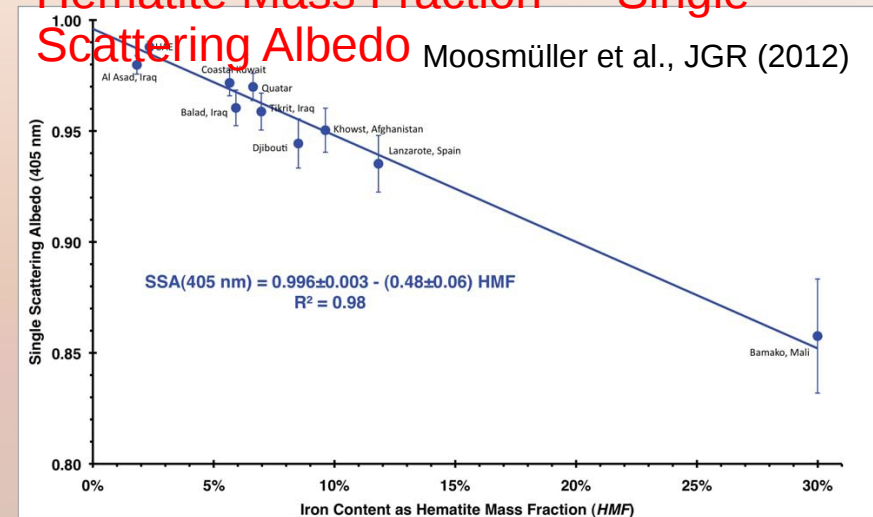
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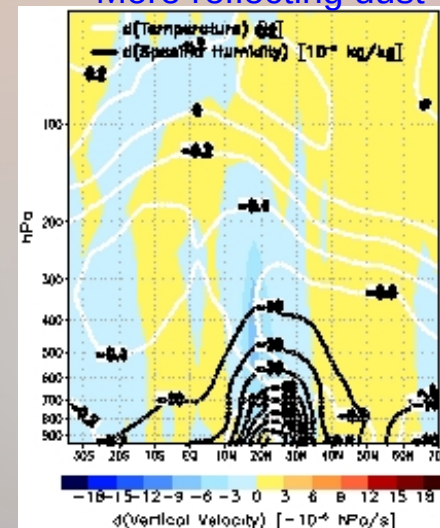
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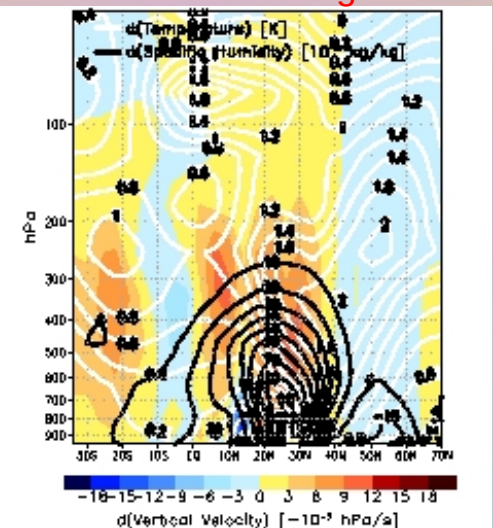


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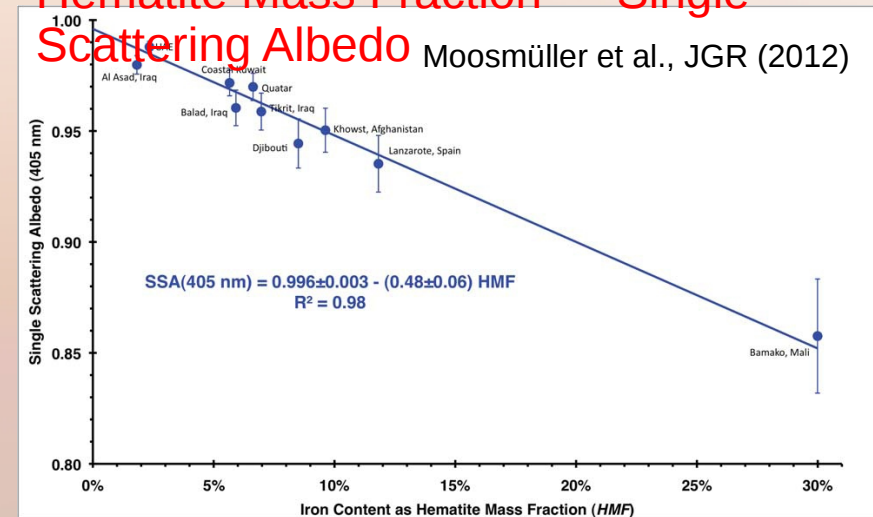
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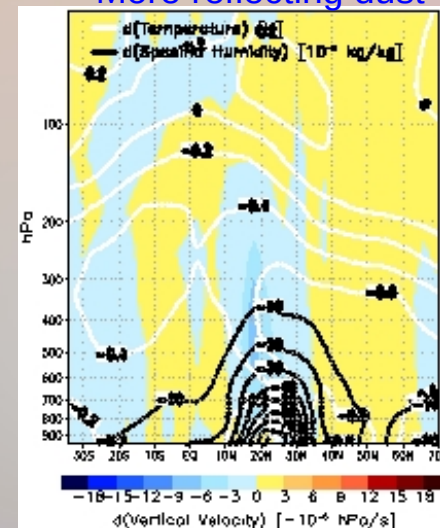
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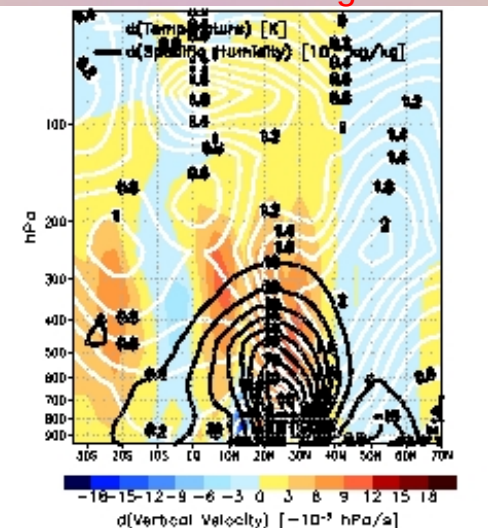


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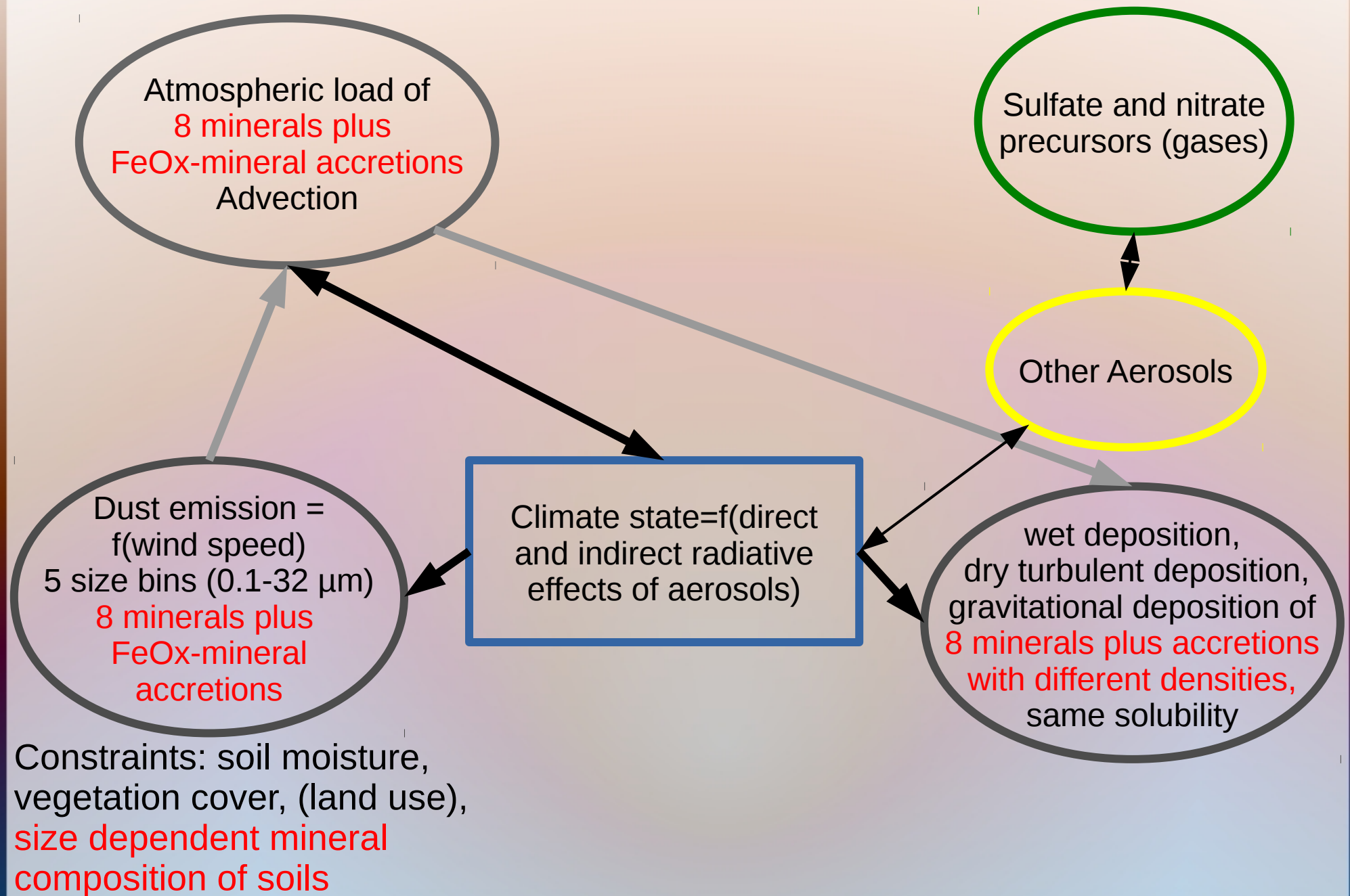


More absorbing dust

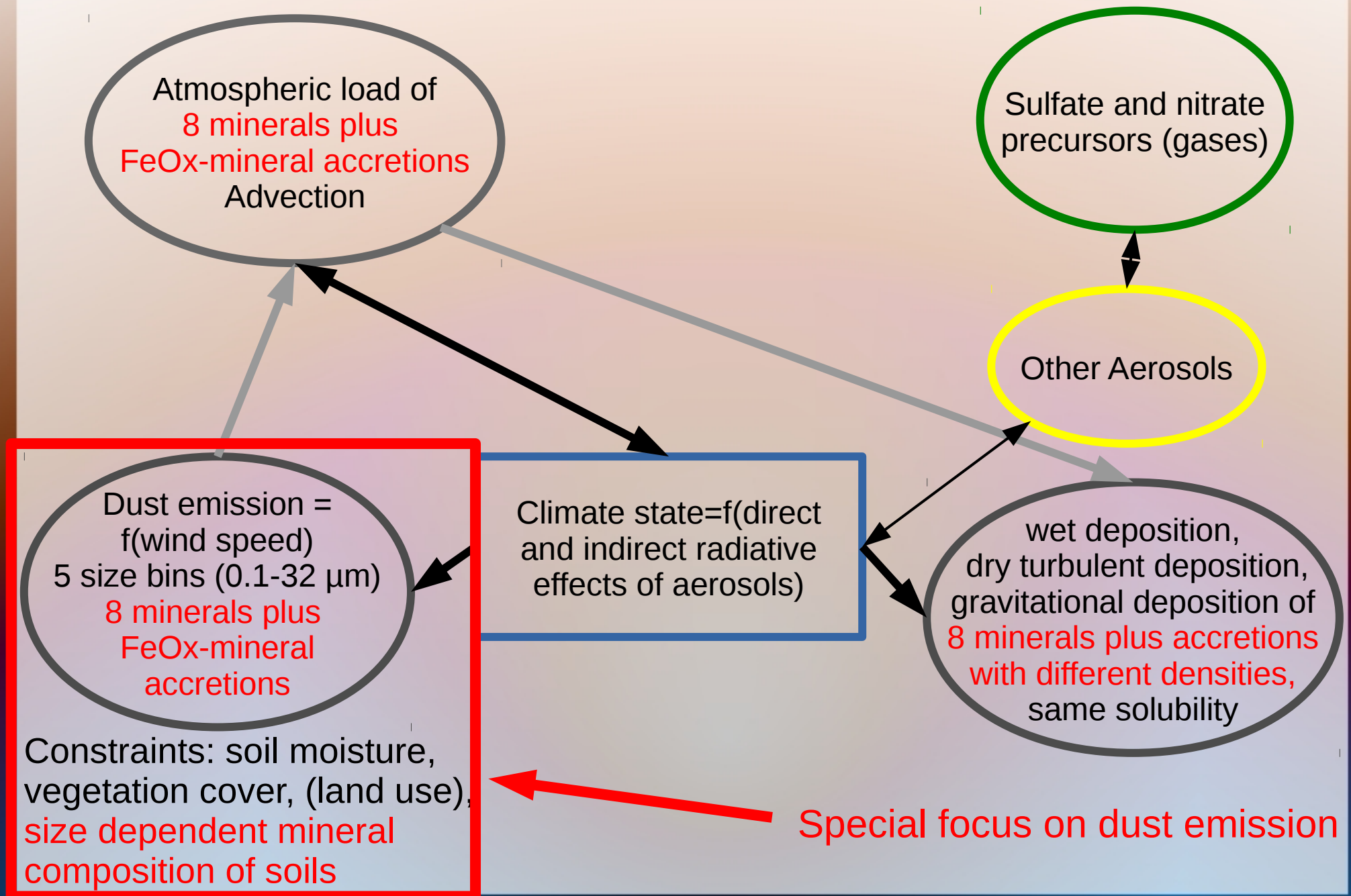


Shaded: Blue: Less Upward Red: More Upward

New Soil Dust in NASA-GISS ModelE2 With Mineralogy



New Soil Dust in NASA-GISS ModelE2 With Mineralogy



The Main Data Sets Needed

1. Mean Mineralogical Table (MMT) by Claquin et al., JGR (1999) + Nickovic et al., ACP (2012)

The mineralogical composition of soils varies with the soil type. The MMT provides this information for 28 arid soil types

CLAQUIN ET AL.: MODELING DUST MINERALOGY

22,247

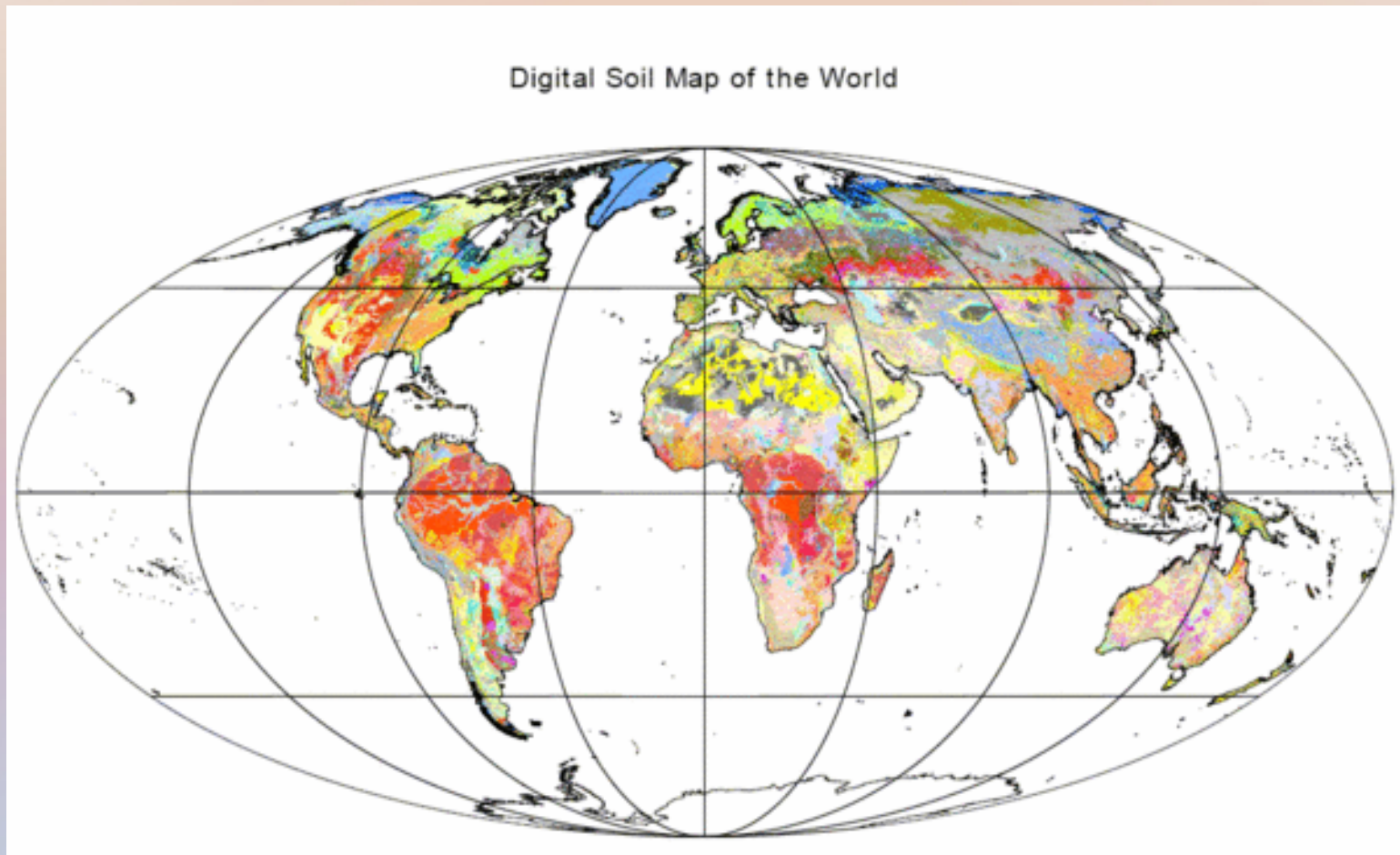
Table 2. Mean Mineralogy Table

| Soil Types | Clay Fraction | | | | | Silt Fraction | | | | | N | A |
|------------|---------------|-----|-----|-----|-----|---------------|-----|-----|-----|-----|----|----|
| | Ill | Kao | Sme | Cal | Qua | Qua | Fel | Cal | Hem | Gyp | | |
| I | | | | | | | | | | | | |
| Average | 40 | 20 | 29 | 4 | 7 | 53 | 40 | 6 | 1 | 1 | 6 | 7 |
| s.d. | 13 | 18 | 17 | 4 | 3 | 14 | 12 | 6 | 1 | 1 | | |
| Jc | | | | | | | | | | | | |
| Average | 22 | 9 | 46 | 11 | 12 | 31 | 39 | 30 | 0 | 2 | 10 | 2 |
| s.d. | 4 | 3 | 13 | 2 | 3 | 2 | 4 | 3 | 0 | 1 | | |
| Je | | | | | | | | | | | | |
| Average | 18 | 23 | 55 | 1 | 3 | 86 | 10 | 2 | 1 | 1 | 12 | 1 |
| s.d. | 10 | 11 | 19 | 1 | 2 | 13 | 5 | 2 | 1 | 1 | | |
| SD | | | | | | | | | | | | |
| Average | 50 | 9 | 26 | 1 | 14 | 92 | 6 | 1 | 1 | 1 | 8 | 12 |
| s.d. | 12 | 5 | 6 | 1 | 4 | 3 | 3 | 1 | 1 | 1 | | |
| ST | | | | | | | | | | | | |
| Average | 39 | 4 | 26 | 29 | 1 | 5 | 1 | 93 | 1 | 26 | 14 | 3 |
| s.d. | 12 | 3 | 18 | 3 | 1 | 4 | 1 | 27 | 1 | 10 | | |

An updated table has just been published by Journet et al. ACP (2014), which is not used here.

2. Digital Soil Map of the World (DSMW) (FAO-UNESCO, 2007)

Geographical distribution of dominant top soil types
(5'x5' latitude by longitude)



3. FAO/STATSGO Soil Texture Fractions

Geographical distribution of clay, silt, and sand fraction for soil texture types
(5'x5' latitude by longitude)

| Class No. | Soil texture class | Sand [%] | Silt [%] | Clay [%] |
|-----------|--------------------|----------|----------|----------|
| 1 | Sand | 92 | 5 | 3 |
| 2 | Loamy Sand | 82 | 12 | 6 |
| 3 | Sandy Loam | 58 | 32 | 10 |
| 4 | Silt Loam | 17 | 70 | 13 |
| 5 | Silt | 10 | 85 | 5 |
| 6 | Loam | 43 | 39 | 18 |
| 7 | Sandy Clay Loam | 58 | 15 | 27 |
| 8 | Silty Clay Loam | 10 | 56 | 34 |
| 9 | Clay Loam | 32 | 34 | 34 |
| 10 | Sandy Clay | 52 | 6 | 42 |
| 11 | Silty Clay | 6 | 47 | 47 |
| 12 | Clay | 22 | 20 | 58 |

Fully dispersed soils!

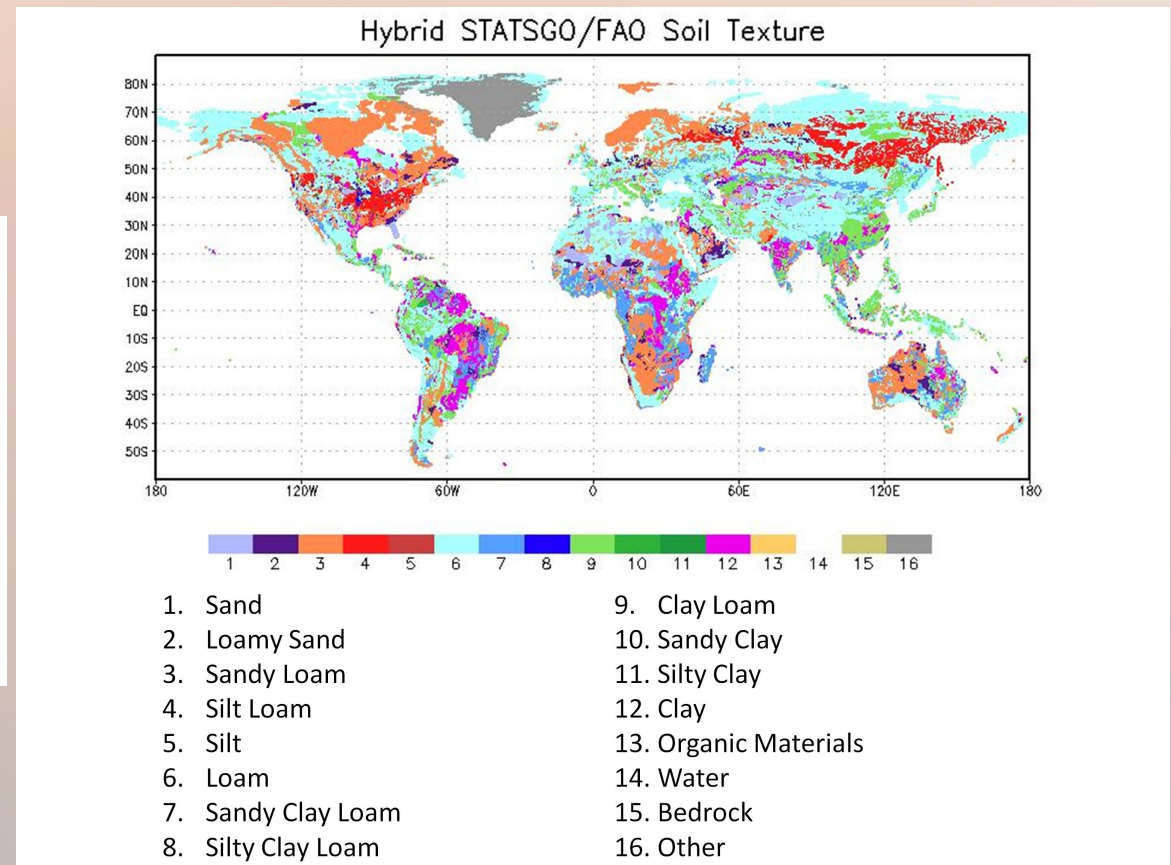
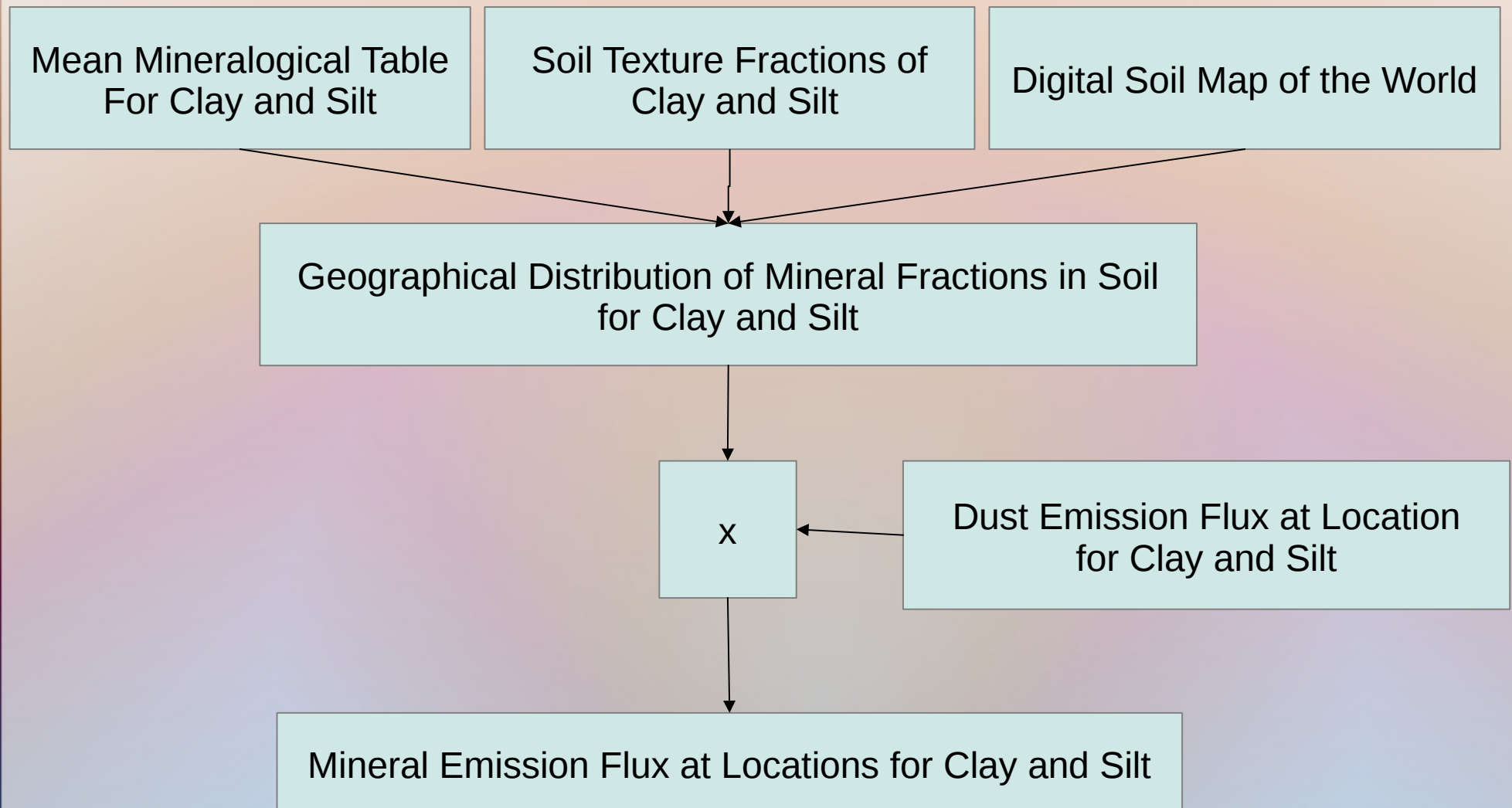


Figure source: <http://ldas.gsfc.nasa.gov/gldas/GLDASsoils.php>

How to Obtain the Emitted Mineral Fractions?

The Simple Approach. Case 1 - Soil Mineral Fraction (SMF) Method



Challenge: Emission of Minerals from Soils

Previous dust models with mineralogy have assumed 1 to 1 translation of mineral fractions in soil data sets to mineral fractions of dust aerosols

Emission from Soils: Aggregated and Fragmented Dust Particles

Shao et al. (2011)

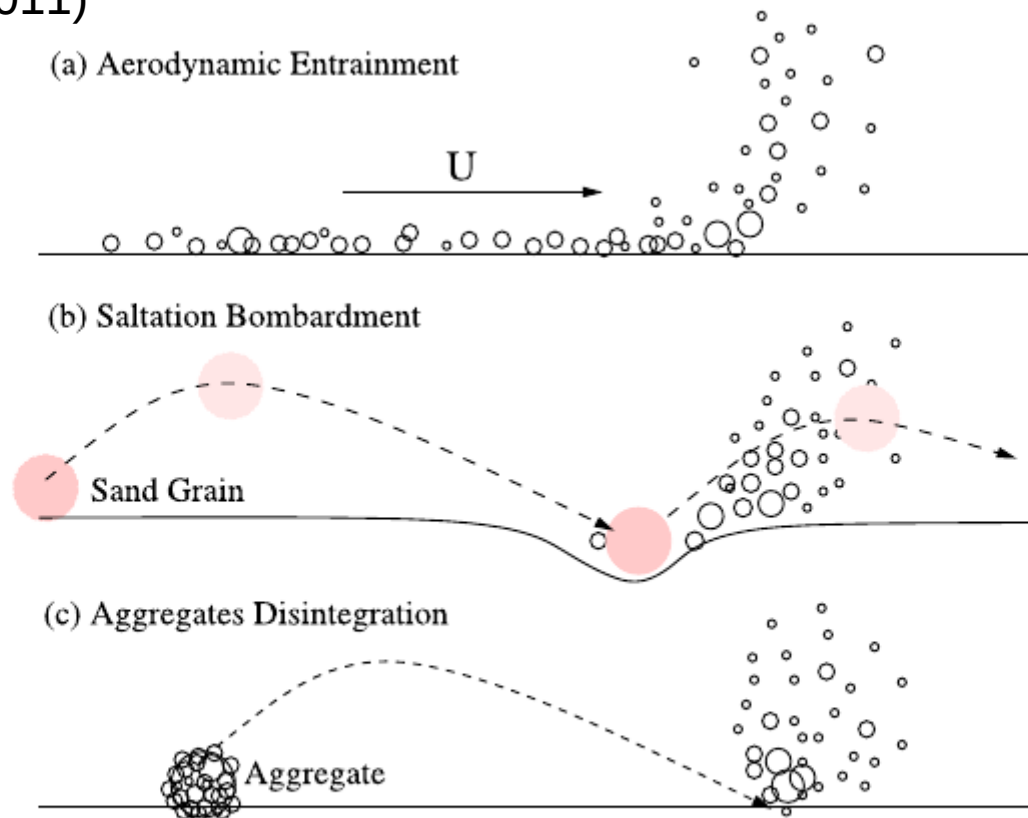


Figure 1. Mechanisms for dust emission. (a) F_a , dust emission by aerodynamic lift; (b) F_b , by saltation bombardment (sand blasting); and (c) F_c , through disaggregation of aggregates or fragments (self-abrasion).

Soil Texture and Mineral Fractions Determined Using Techniques Leading to Nearly Full Destruction of Aggregates

Figure 5-5. Laboratory sieves for mechanical analysis of grain size distribution. Shown (right to left) are sieve Nos. 3/8-in. (9.5-mm), No. 10 (2.0-mm), No. 40 (250- μ m) and No. 200 (750- μ m) and example soil particle sizes including (right to left): medium gravel, fine gravel, medium-coarse sand, silt, and dry clay (kaolin).

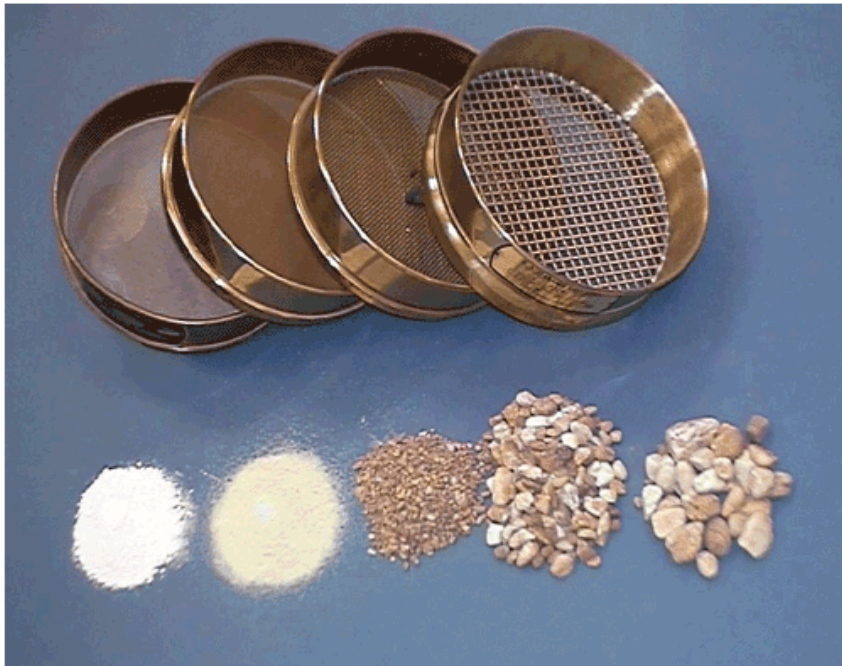


Figure 5-6. Soil hydrometer apparatus (<http://www.ce.siu.edu/>).



Source: <http://www.fhwa.dot.gov/engineering/geotech/pubs/05037/05a.cfm>



Wet Sieved Soil Texture Fractions \neq Size Distribution of Eroded Soils

Wet Sieved Soil Texture Fractions \neq Suspended Dust Size Distribution

Wet Sieved Clay/Silt Mineral Fractions \neq Mineral Fractions of Suspended Dust

Mineral Fractions in Dust vs. Mean Mineralogical Table

Abundance of mineral groups over particle size at Tinfou, Morocco:

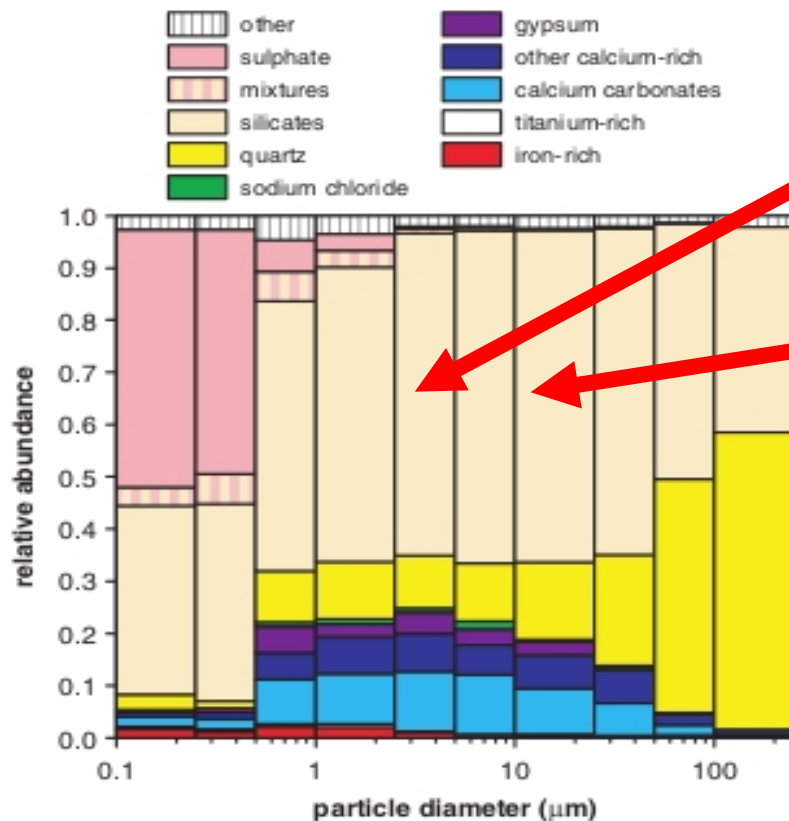


Fig. 11. Average relative volume abundance of the different particle classes at Tinfou ground station. Kandler et al., Tellus B (2009)

- Significant Fractions of so called “Clays” like illite and kaolinite are being found in silt size range, but missing in silt size range of MMT
- Silicates can even have the largest fraction in silt size range.
- Quartz may be dominant only for largest sizes.
- Feldspar, gypsum missing in clay size range of MMT

Another challenge with MMT:
No accretions of iron oxides! Iron oxides have higher density than other minerals and fall out faster. How are they transported in the model?

Mineral Fractions in Dust vs. Mean Mineralogical Table

Abundance of mineral groups over particle size at Tinfou, Morocco:

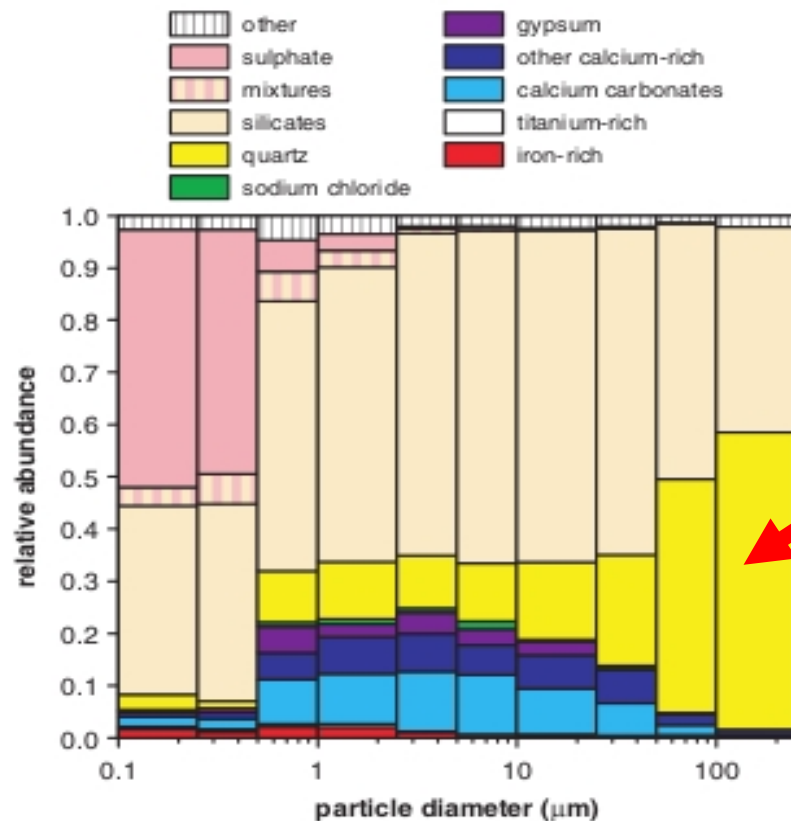


Fig. 11. Average relative volume abundance of the different particle classes at Tinfou ground station. Kandler et al., Tellus B (2009)

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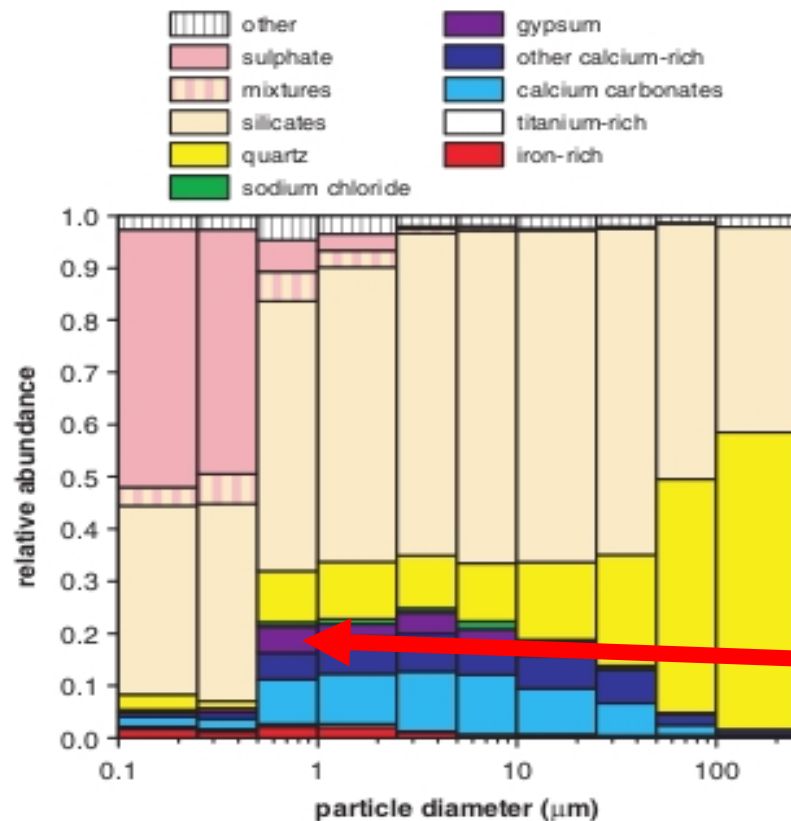


Fig. 11. Average relative volume abundance of the different particle classes at Tinfou ground station. Kandler et al., Tellus B (2009)

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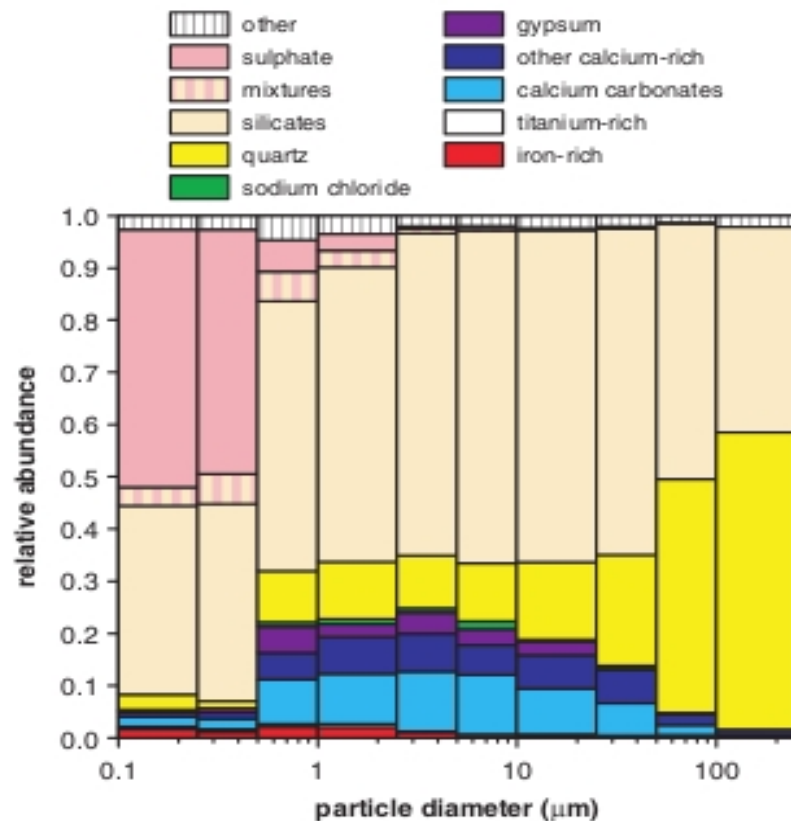


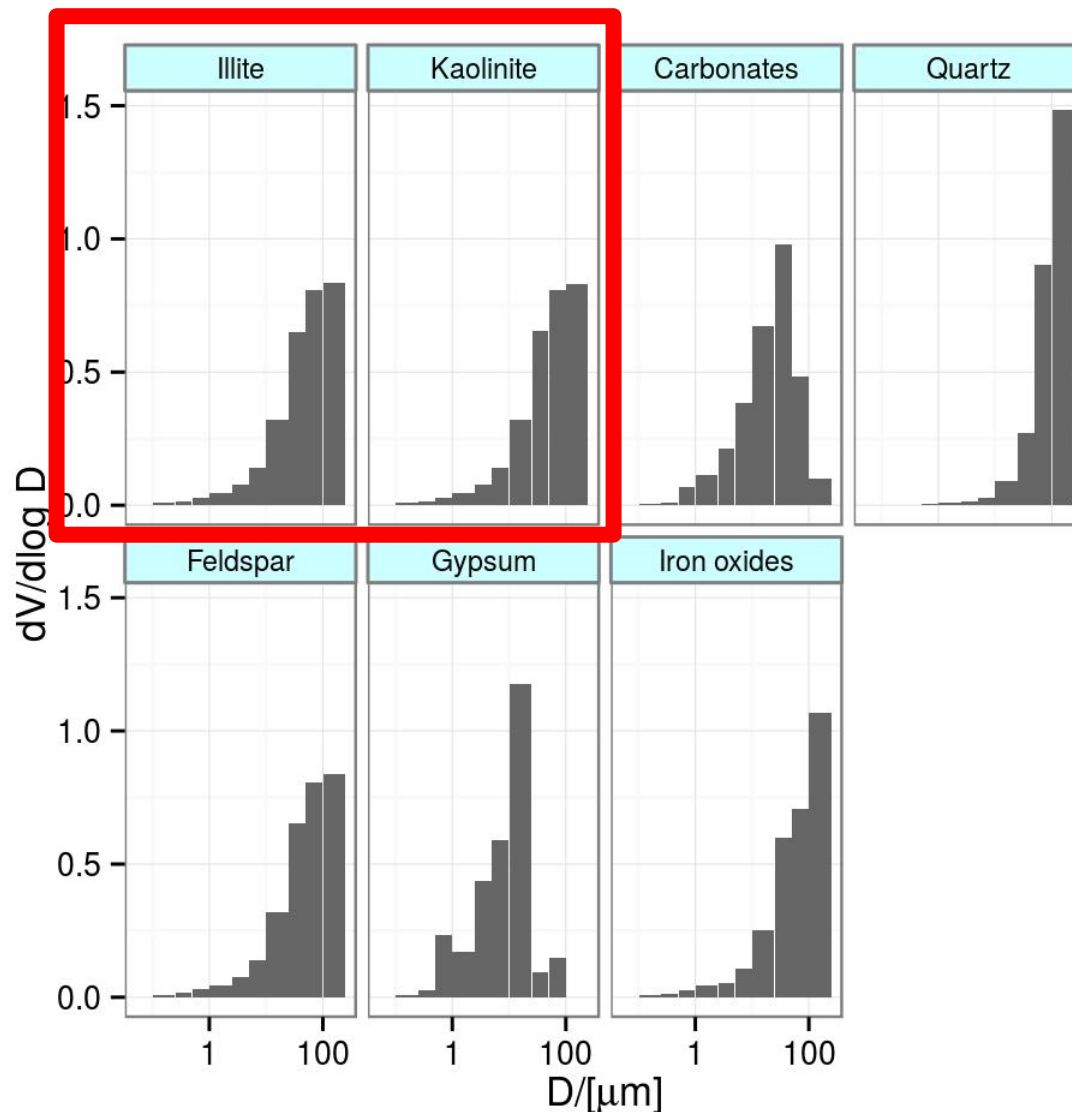
Fig. 11. Average relative volume abundance of the different particle classes at Tinfou ground station.

Kandler et al., Tellus B (2009)

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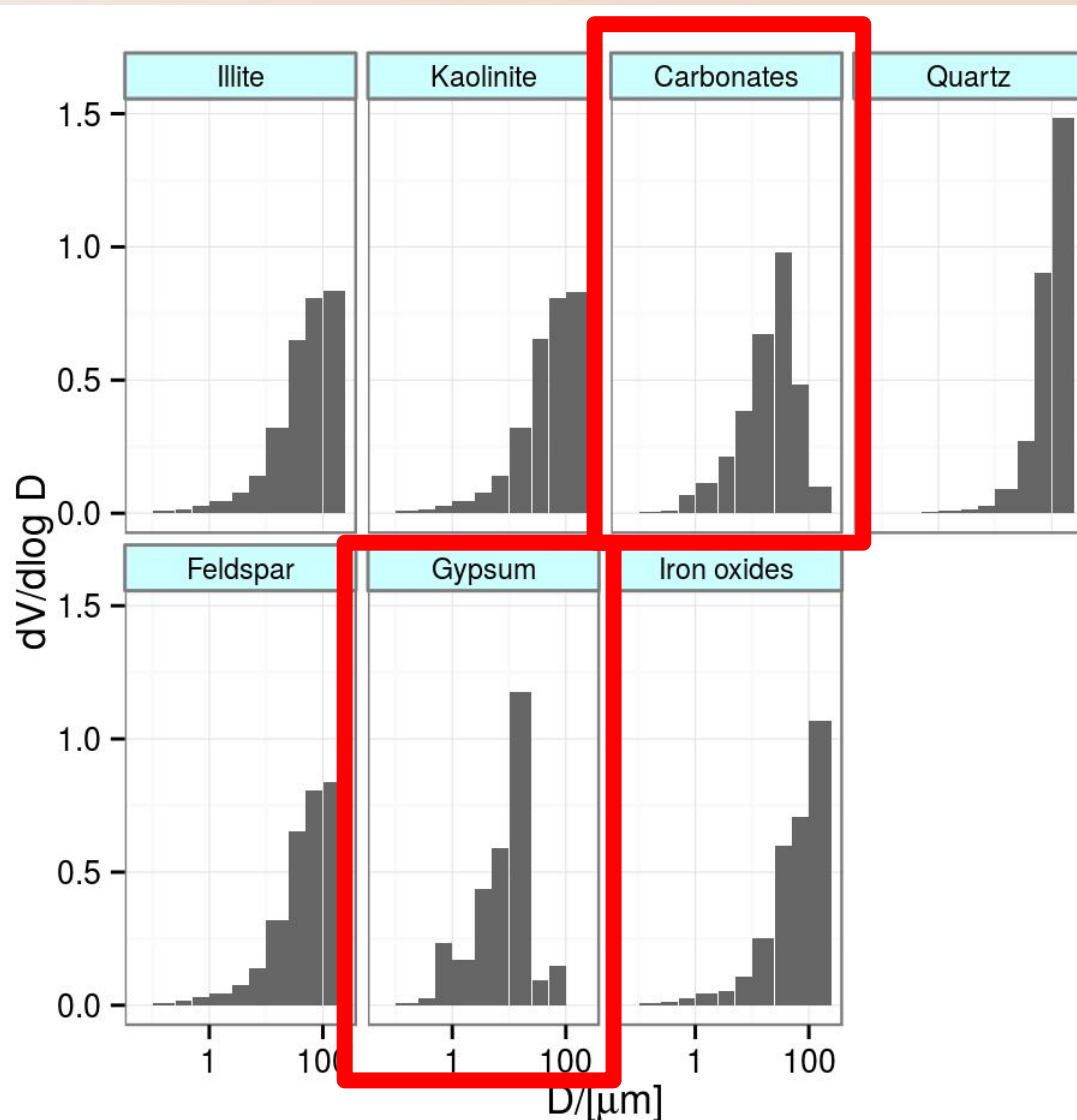
Another challenge with MMT:
No accretions of iron oxides! Iron oxides have higher density than other minerals and fall out faster. How are they transported in the model?

Normalized Volume Size Distribution of Mineral Fractions in Dust Derived From Data Provided by Kandler et al. Tellus B (2009)



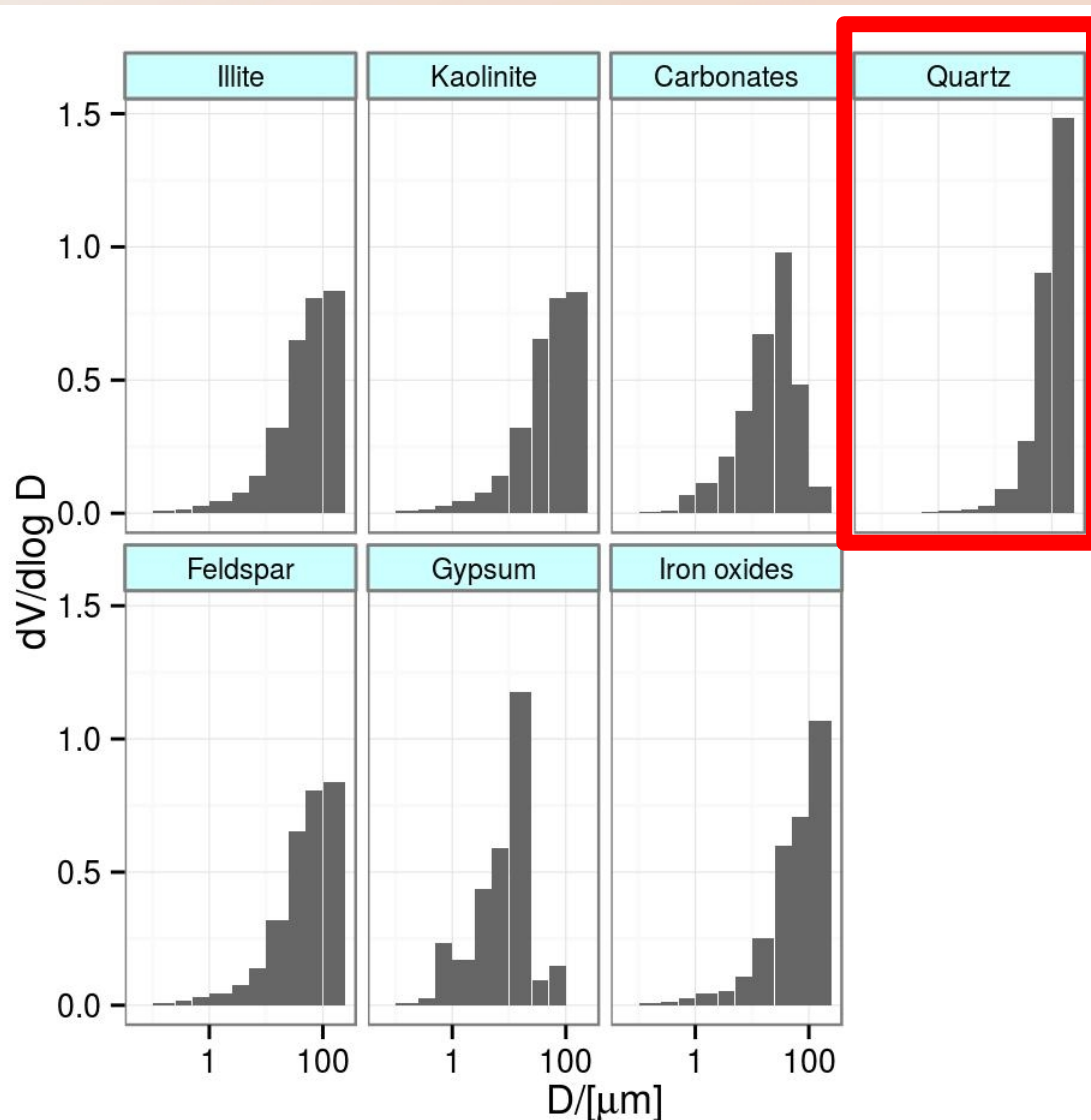
- Illite and kaolinite: Similar volume size distribution; most of the volume (mass) is found in higher particle size classes, even beyond silt size range (probably mostly due to aggregation)
- The carbonates and gypsum peak in the coarse silt size class
- Distinctive size distribution of quartz with steep increase in the volume distribution for largest particle sizes

Normalized Volume Size Distribution of Mineral Fractions in Dust Derived From Data Provided by Kandler et al. Tellus B (2009)



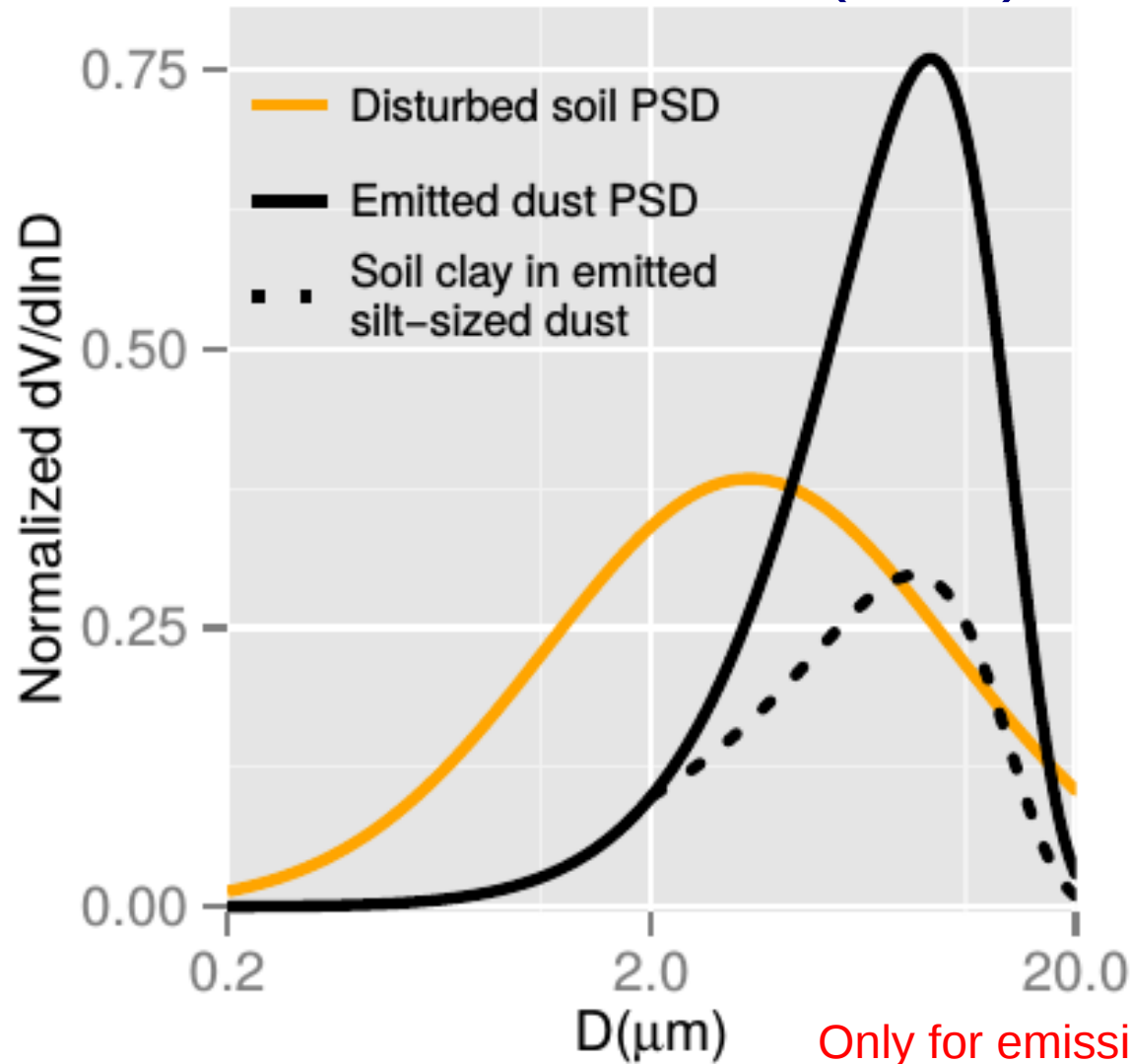
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Aerosol Mineral Fraction (AMF) Method

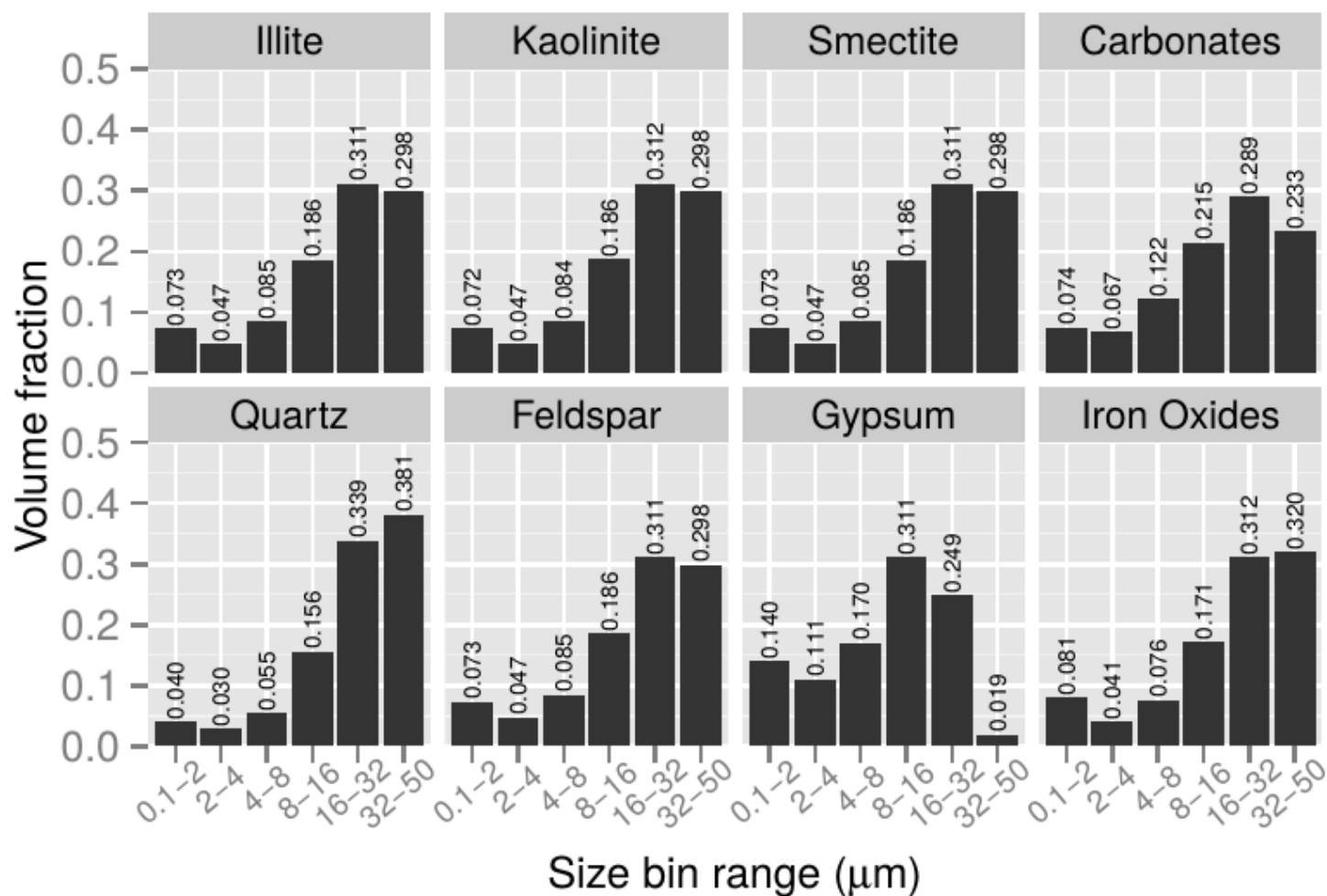


Fully disbursed mineral mass in soil

Re-aggregation

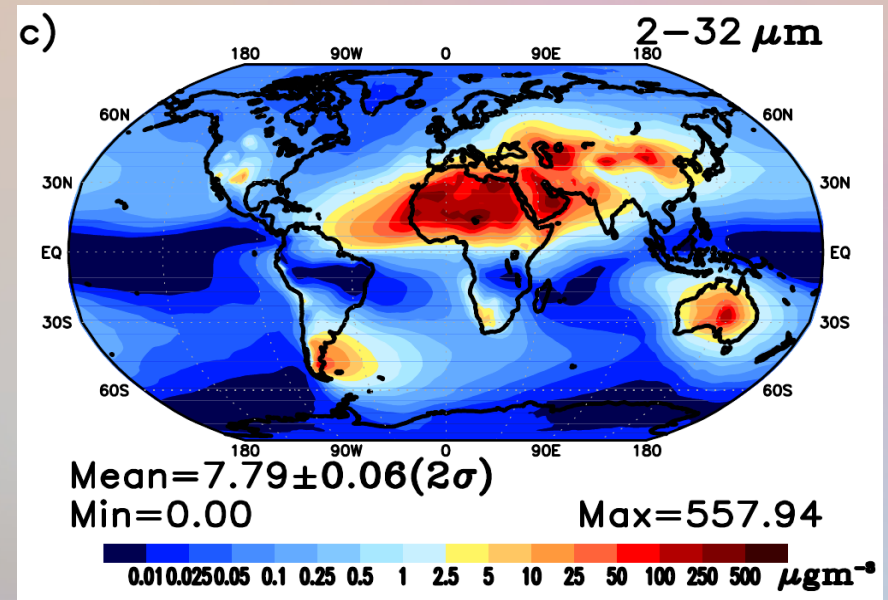
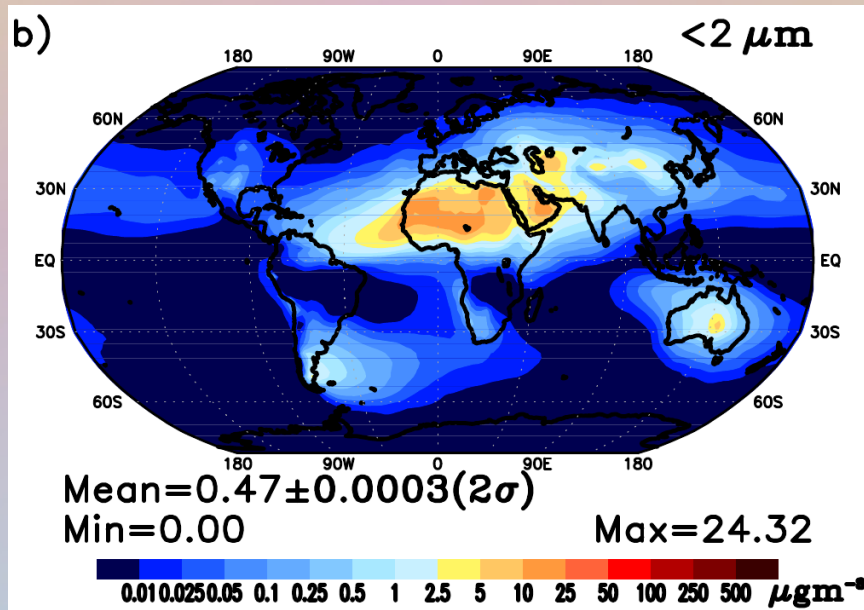
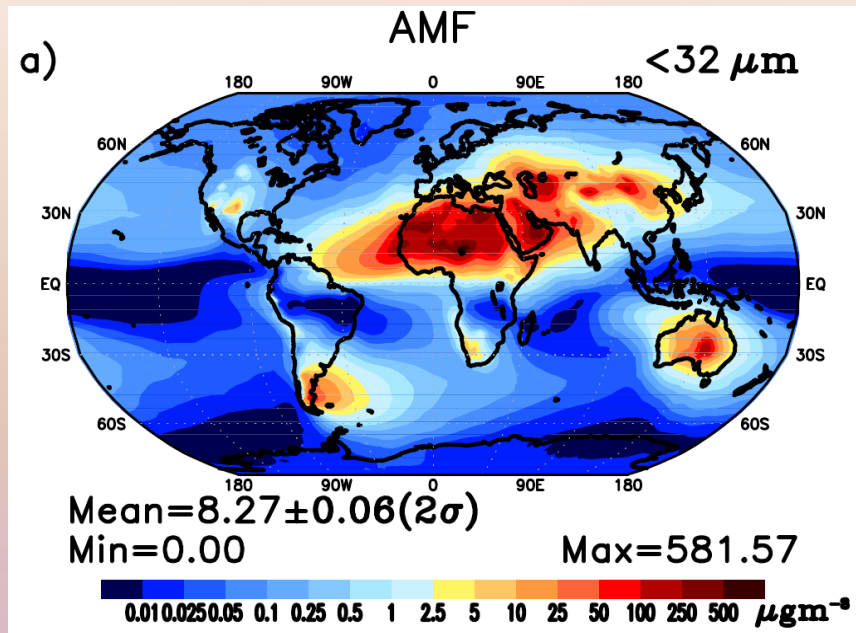
Reconstructed size-dependent mineral mass in aerosols

Volume Size Distribution of Emitted Minerals in ModelE2



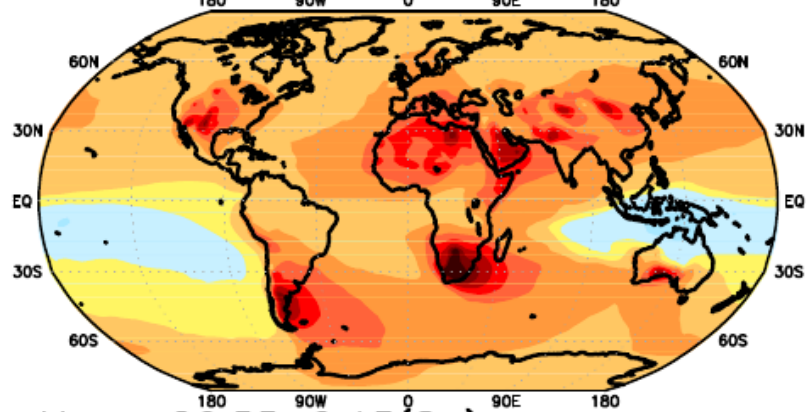
Perlwitz et al., ACPD (2015a)

Surface Concentration of Total Dust



Simulated Mineral Fractions of Dust Column Mass

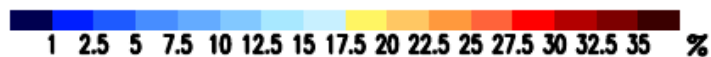
a) Quartz $<32\ \mu\text{m}$



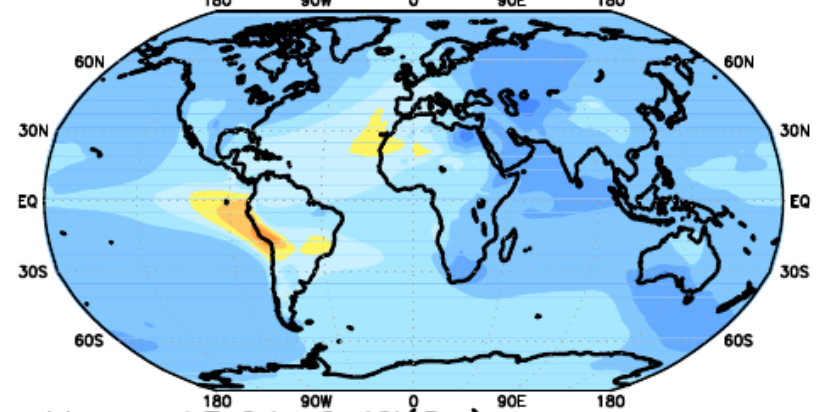
Mean= $26.55 \pm 0.15(2\sigma)$

Min=14.26

Max=39.34



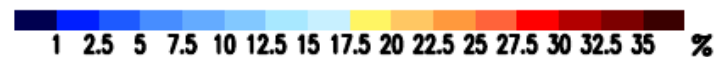
b) Feldspar $<32\ \mu\text{m}$



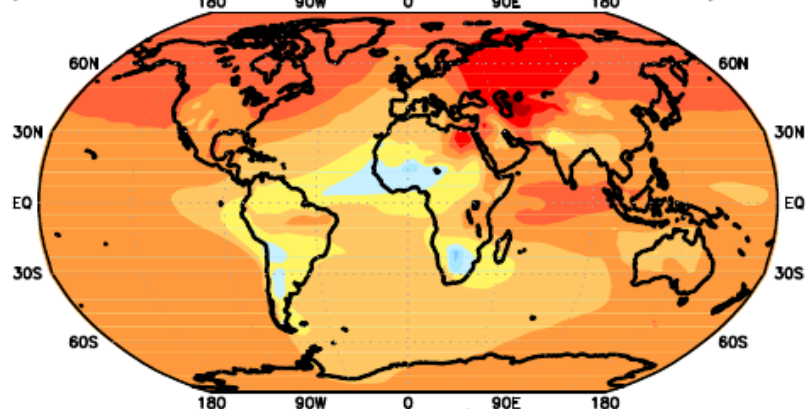
Mean= $13.01 \pm 0.47(2\sigma)$

Min=5.77

Max=24.18



c) Illite $<32\ \mu\text{m}$



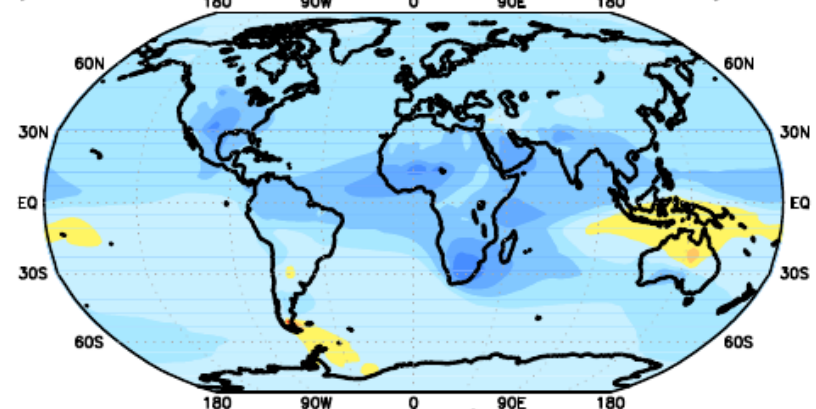
Mean= $21.33 \pm 0.54(2\sigma)$

Min=11.20

Max=33.01



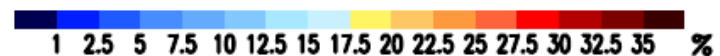
d) Smectite $<32\ \mu\text{m}$



Mean= $11.67 \pm 0.32(2\sigma)$

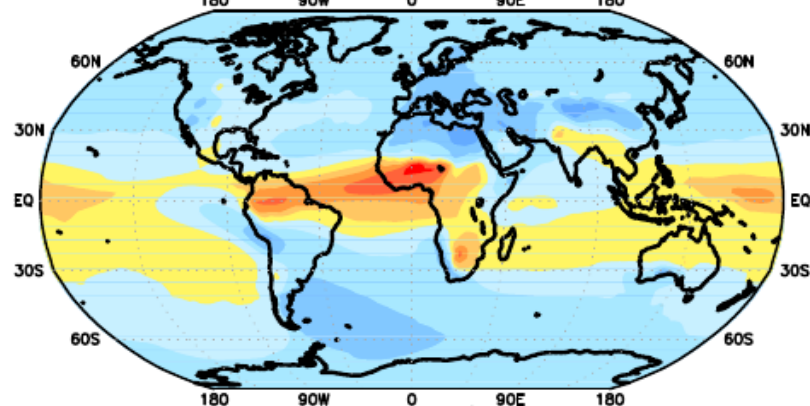
Min=5.01

Max=28.49



Simulated Mineral Fractions of Dust Column Mass

e) Kaolinite <32 μm



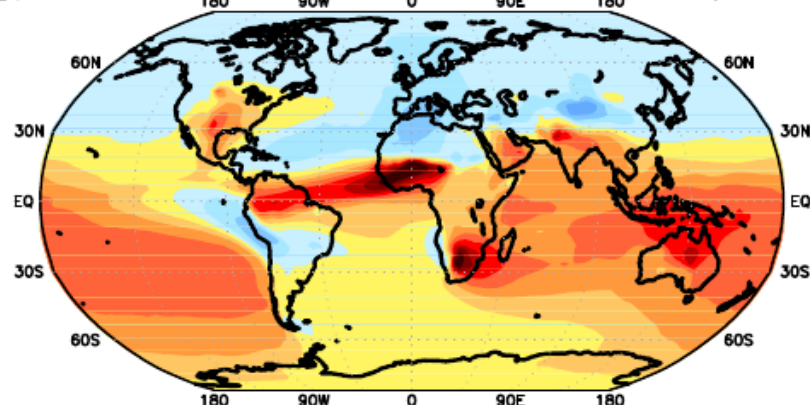
Mean = $16.28 \pm 0.94 (2\sigma)$

Min = 7.95

Max = 29.12



g) Iron Oxides <32 μm



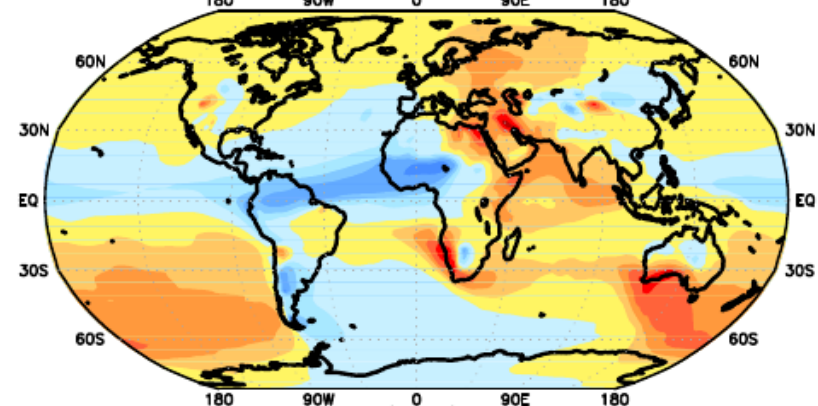
Mean = $1.88 \pm 0.07 (2\sigma)$

Min = 0.83

Max = 3.91



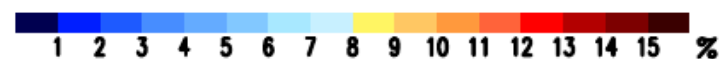
f) Carbonates <32 μm



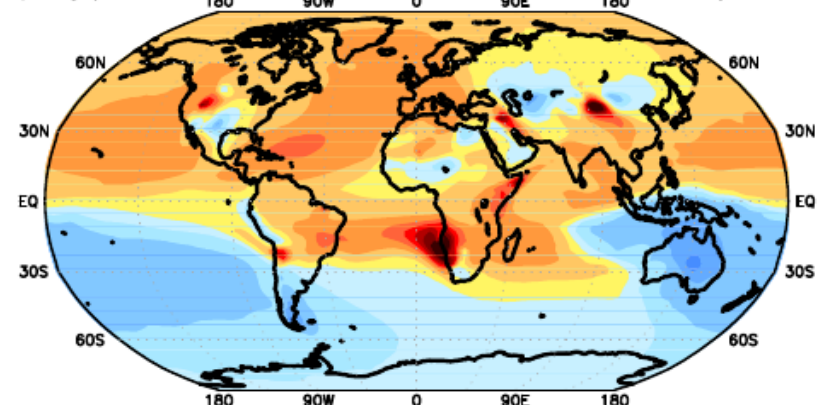
Mean = $7.60 \pm 0.48 (2\sigma)$

Min = 3.30

Max = 13.56



h) Gypsum <32 μm



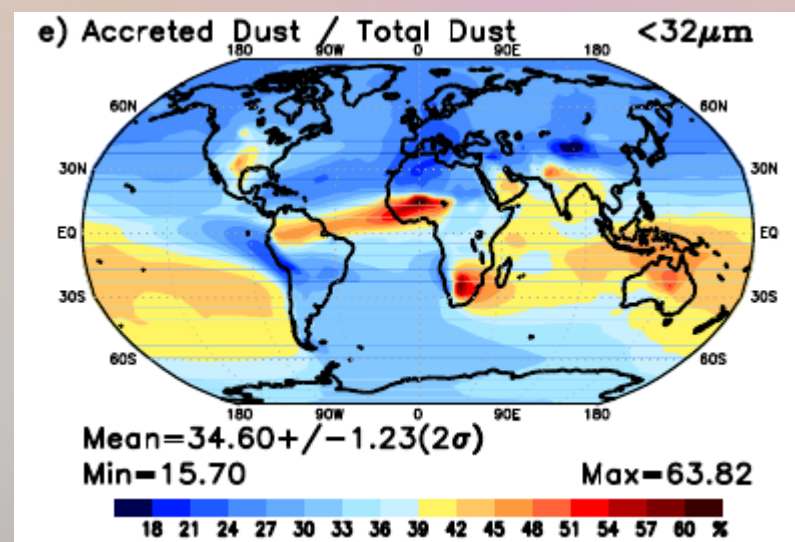
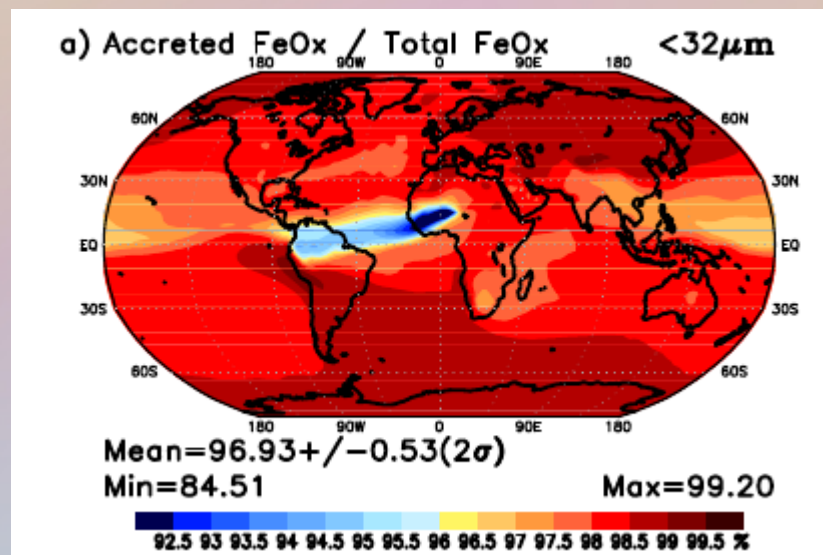
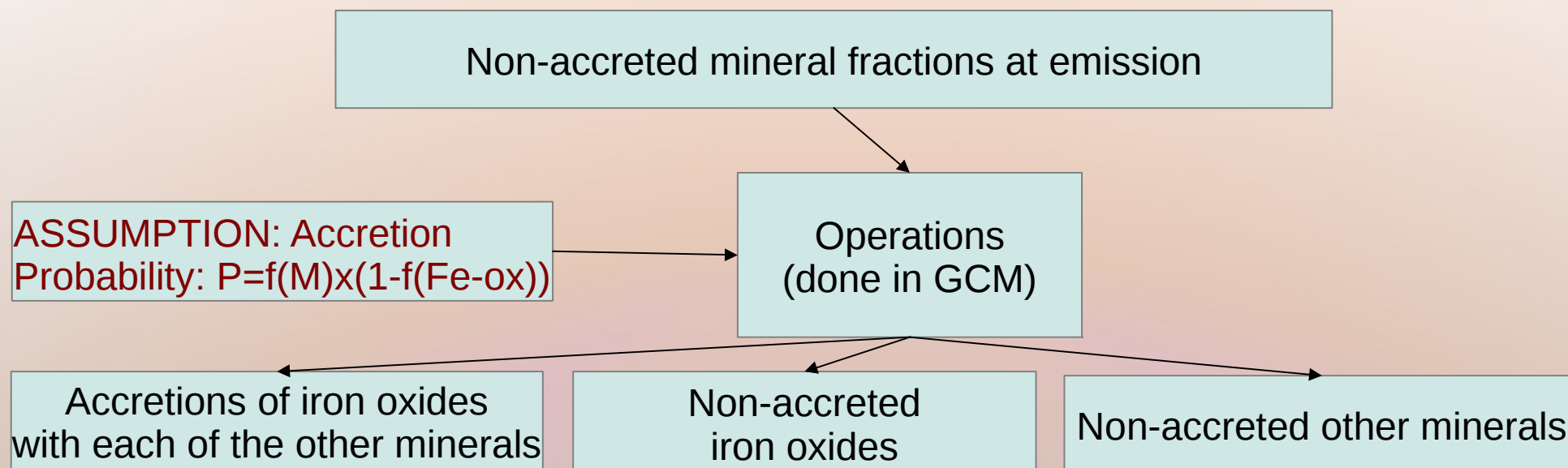
Mean = $1.68 \pm 0.03 (2\sigma)$

Min = 0.66

Max = 3.68



Iron Oxide Accretions With Other Minerals



Almost all of iron oxide mass accreted

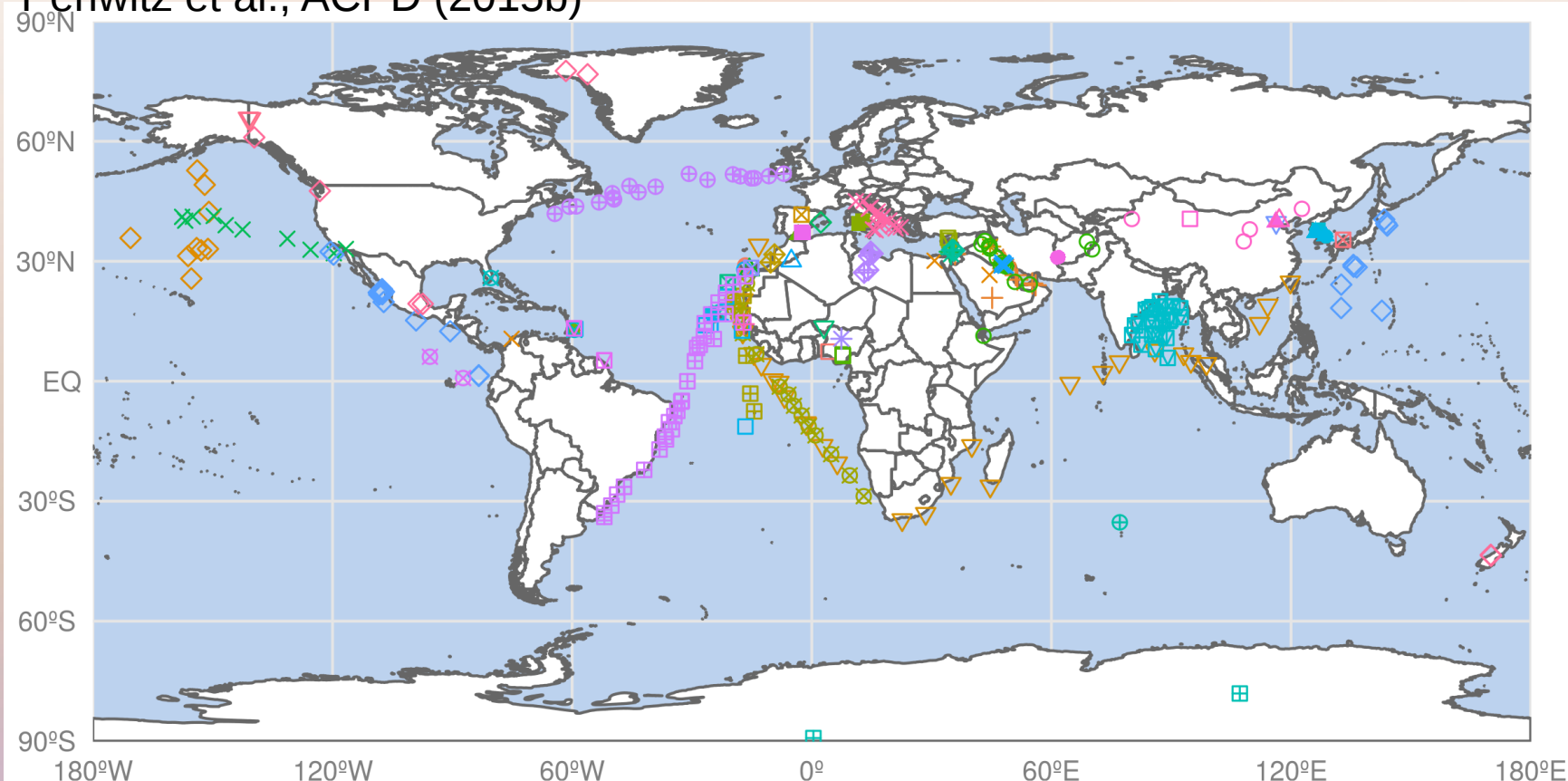
About 35% of non-FeOx dust with FeOx impurities

Evaluation

- Compilation of about 60 references from literature with mineral fraction measurements (Perlwitz et al., ACPD 2015b)
- Limitations: Mostly campaign data or cruises, small sampling size, possible biases depending on the methods; How to compare to model climatology?

Locations of Measurements from Literature

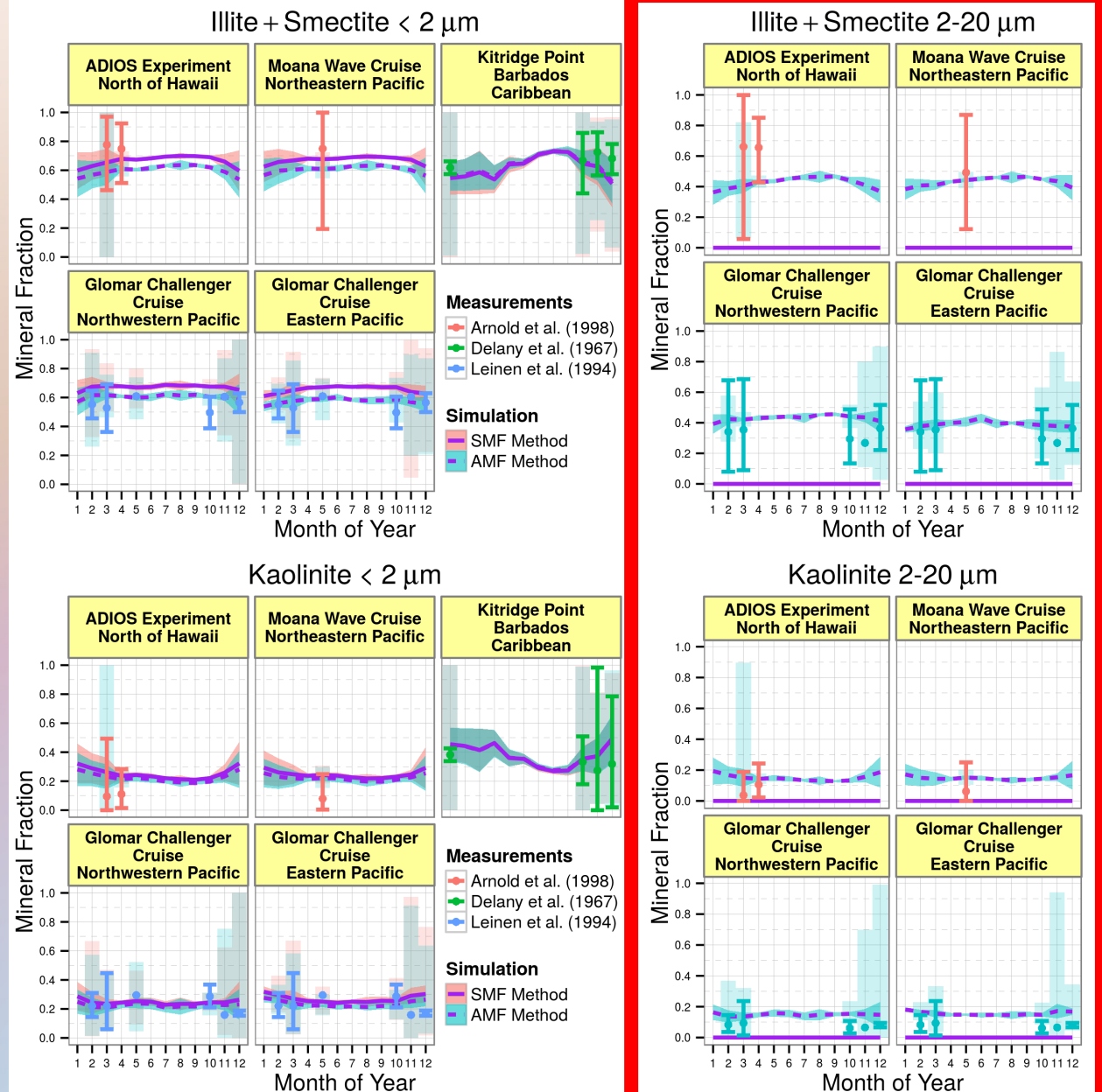
Perlwitz et al., ACPD (2015b)



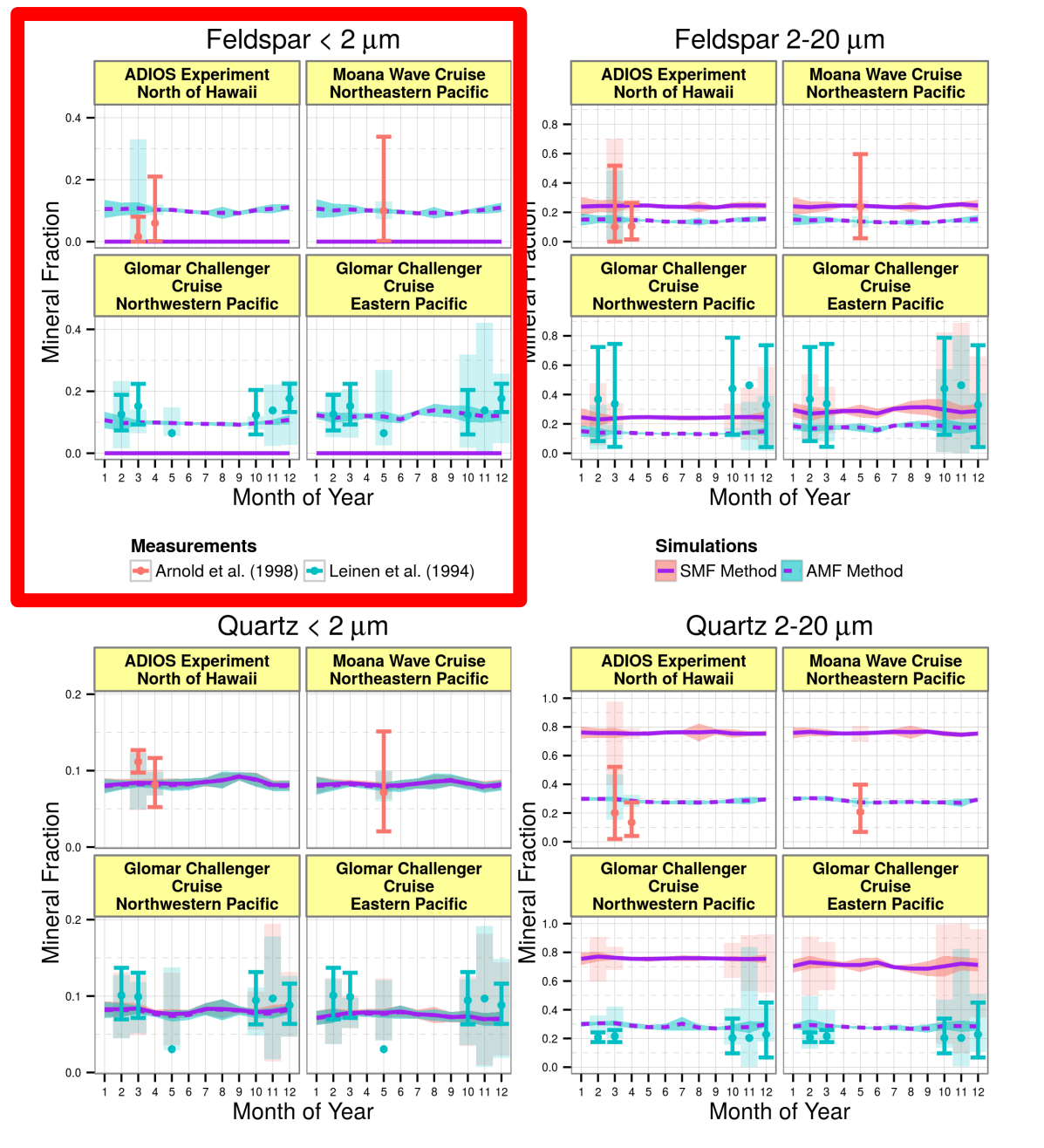
Reference

- | | | | |
|--------------------------------------|---|---------------------------------------|--|
| Adedokun et al. (1989):7.29°N,4.34°E | Delany et al. (1967):13.17°N,59.42°W | Goldberg and Griffin (1970) | Parkin et al. (1972) |
| Alastuey et al. (2005) | Díaz-Hernández et al. (2011):37.17°N,3.52°W | Jeong (2008):37.5°N,126.92°E | Prospero and Bonatti (1969) |
| Al-Awadhi and AlShuaibi (2013) | Enete et al. (2012) | Jeong and Achterberg (2014) | Prospero et al. (1981) |
| Al-Dousari and Al-Awadhi (2012) | Engelbrecht et al. (2009) | Jeong et al. (2014) | Queralt-Mitjans et al. (1993) |
| Al-Dousari et al. (2013) | Engelbrecht et al. (2014):28.07°N,15.45°W | Johnson (1976) | Rashki et al. (2013) |
| Arnold et al. (1998) | Falkovich et al. (2001):32.08°N,34.8°E | Kandler et al. (2007):28.32°N,16.5°W | Shao et al. (2008):39.99°N,116.34°E |
| Aston et al. (1973) | Ferguson et al. (1970) | Kandler et al. (2009):30.24°N,5.6°W | Shen et al., (2006):40.5°N,94.82°E |
| Avila et al. (1997):41.77°N,2.35°W | Fiol et al. (2005):39.63°N,2.65°E | Kandler et al. (2011):14.94°N,23.48°W | Shen et al. (2009) |
| Awadh (2012):33.33°N,44.43°E | Formenti et al. (2008):13.5°N,2.6°E | Khalaf et al. (1985) | Shi et al. (2005):40°N,116.77°E |
| Chester and Johnson (1971a) | Game (1964):25.07°N,20.73°W | Leinen et al. (1994) | Skonieczny et al. (2013):14.41°N,16.96°W |
| Chester and Johnson (1971b) | Ganor (1991):31.78°N,35.22°E | Lu et al. (2006):40°N,116.77°E | Tomadin et al. (1984) |
| Chester et al. (1971) | Ganor et al. (2000) | Menendez et al. (2007):27.97°N,15.6°W | Windom (1969) |
| Chester et al. (1972) | Gaudichet et al. (1989):34.78°S,77.52°E | Moberg et al. (1991):11.07°N,7.7°E | Zdanowicz et al. (2006) |
| Chester et al. (1977) | Gaudichet et al. (1992) | O'Hara et al. (2006) | Zhou and Tazaki (1996):35.48°N,133.07°E |
| Chester et al. (1984) | Glaccum and Prospero (1980) | Parkin et al. (1970) | |

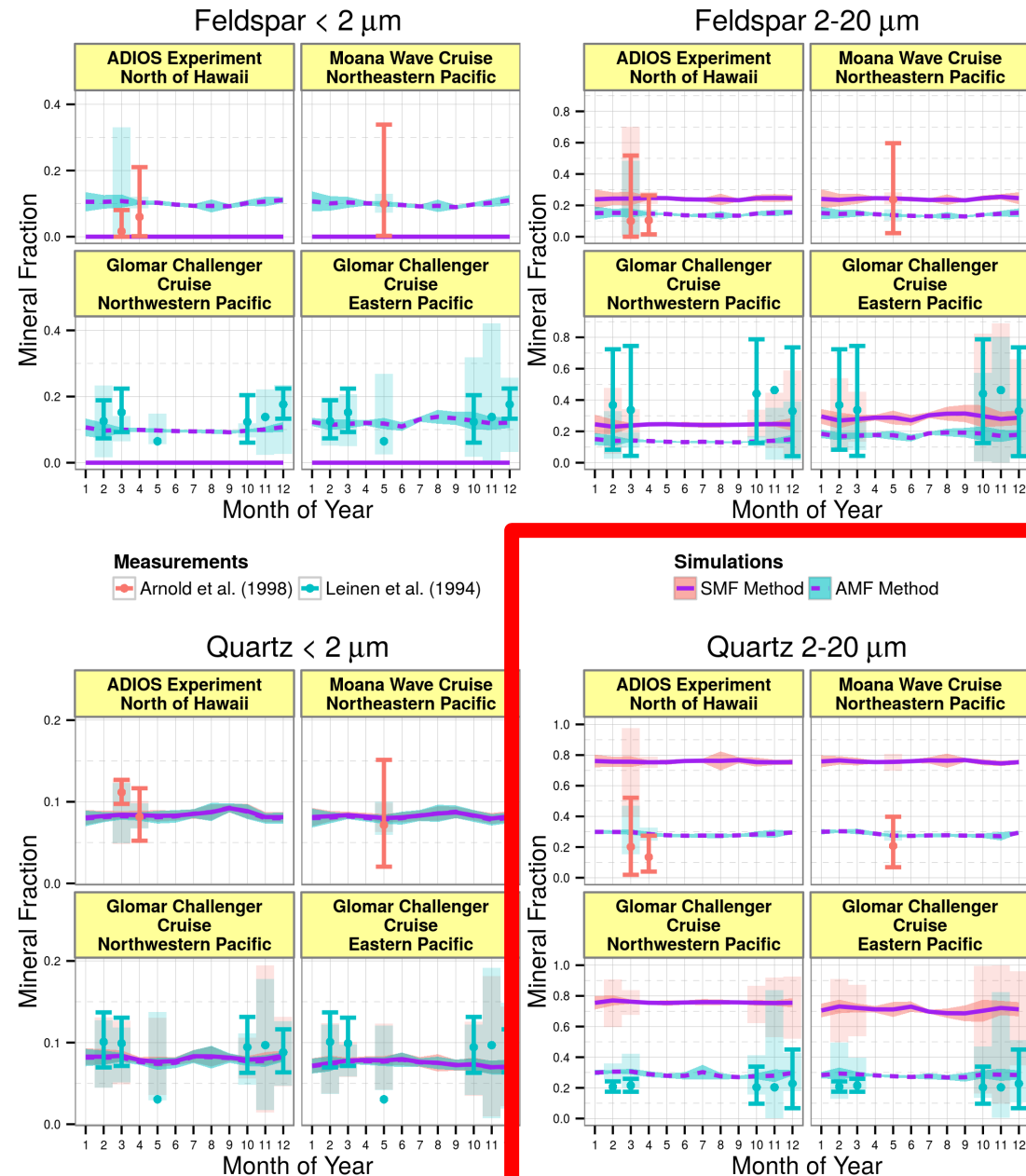
Mineral Fractions Simple (SMF) vs. New (AMF) Method



Mineral Fractions Simple (SMF) vs. New (AMF) Method

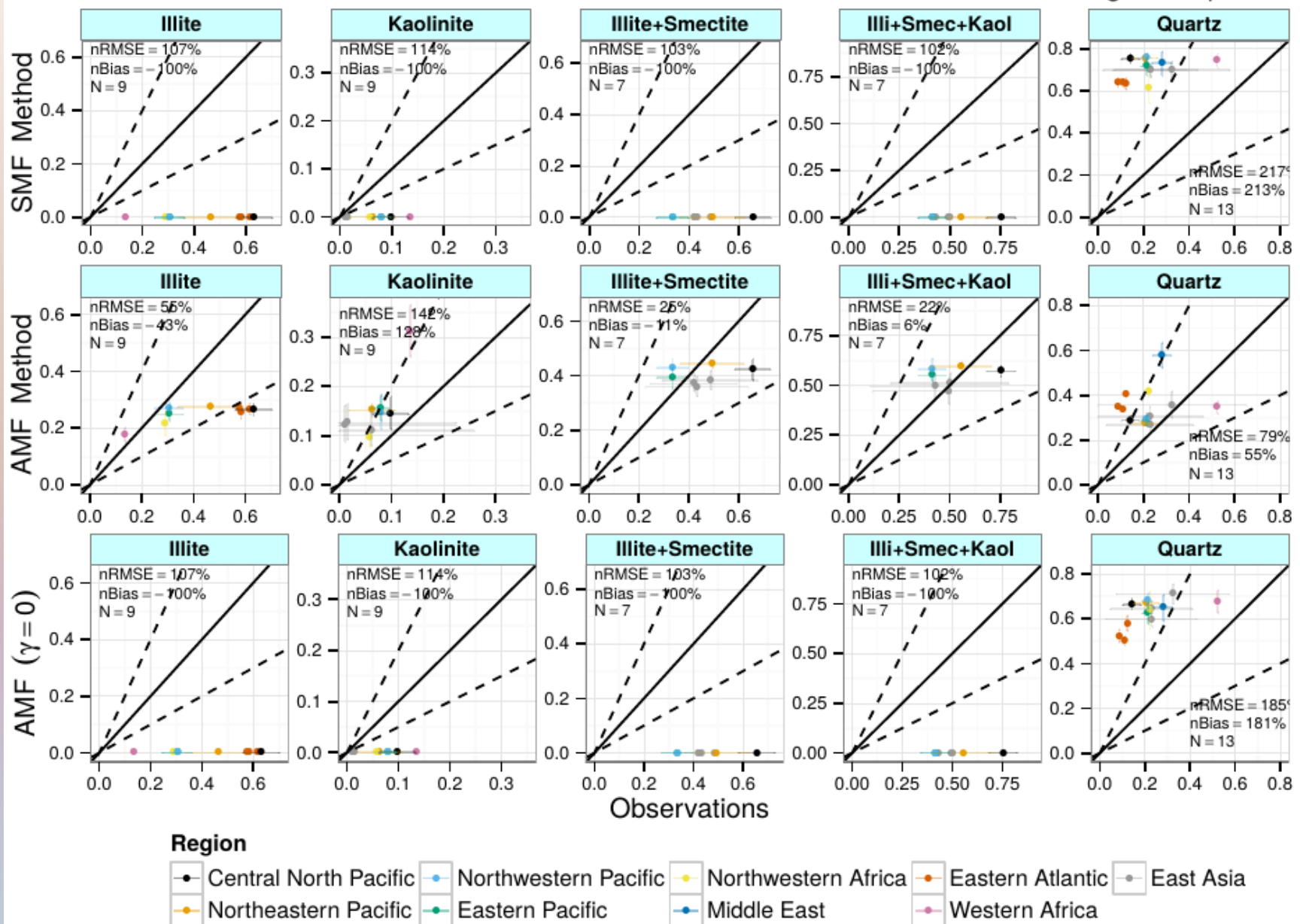


Mineral Fractions Simple (SMF) vs. New (AMF) Method



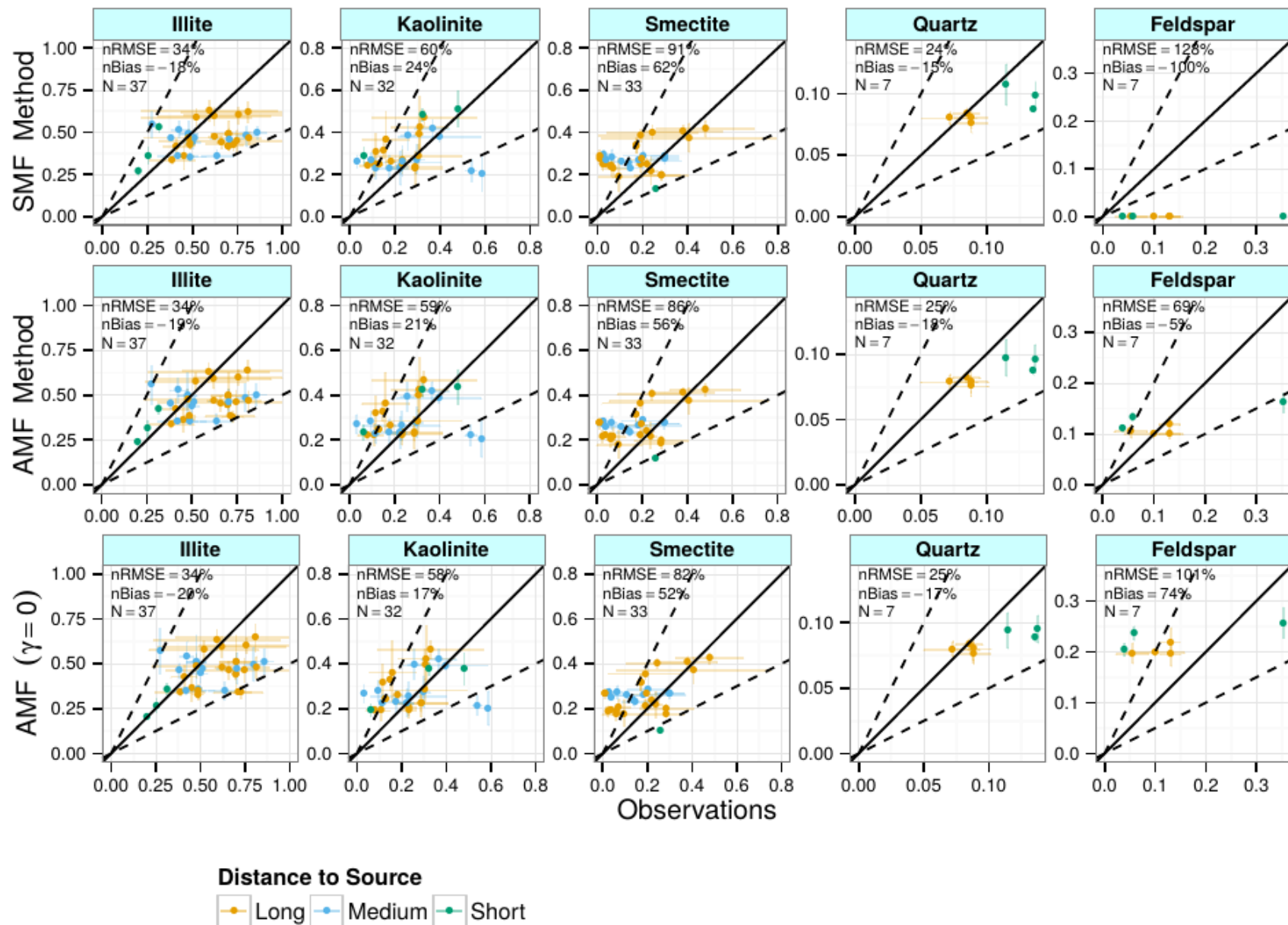
Evaluation - Silt Sized Dust

Fractions of Minerals in Dust – Simulations versus Observations – Size Range: $> 2 \mu\text{m}$



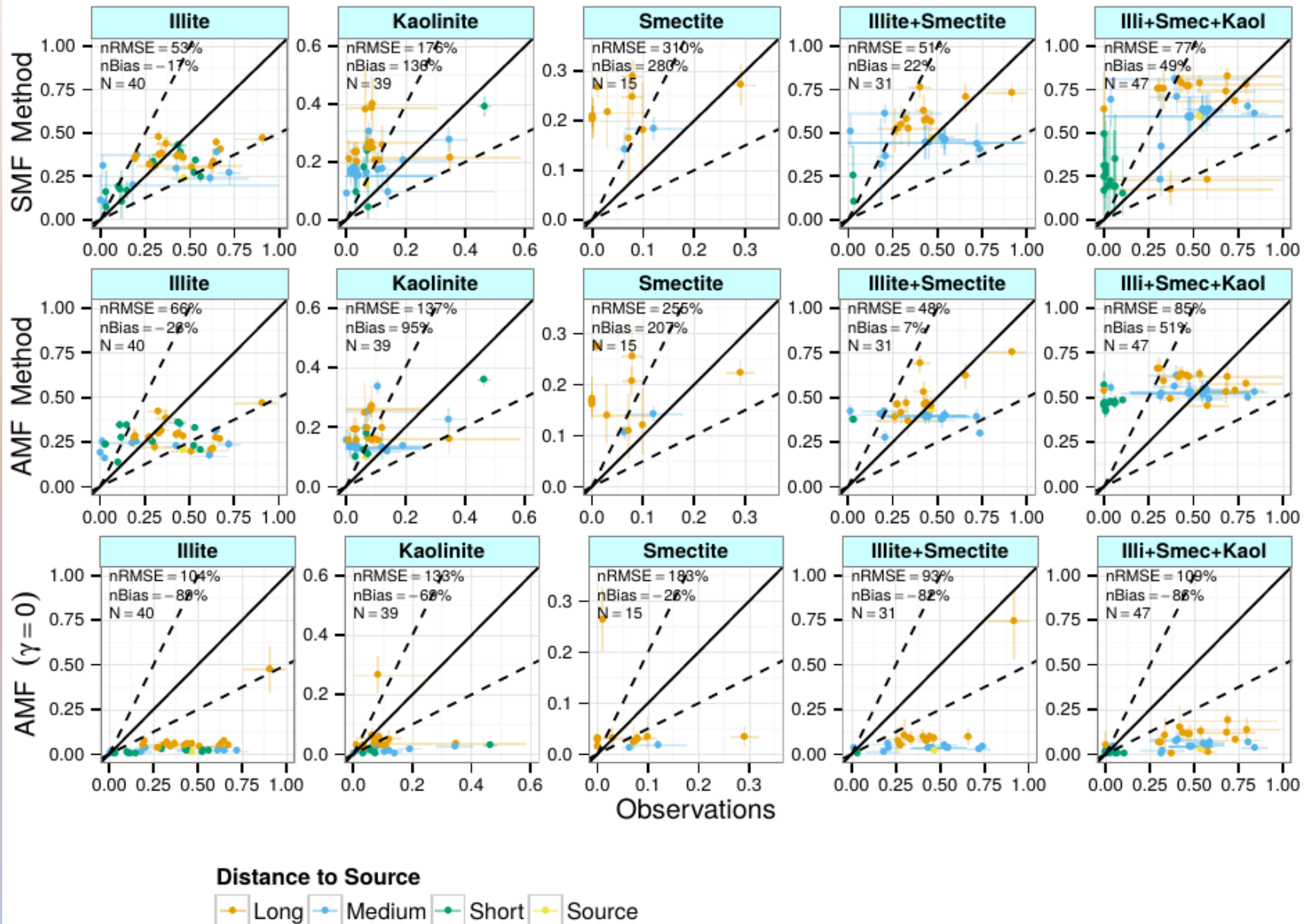
Evaluation - Clay Sized Dust

Fractions of Minerals in Dust – Simulations versus Observations – Size Range: $< 2 \mu\text{m}$



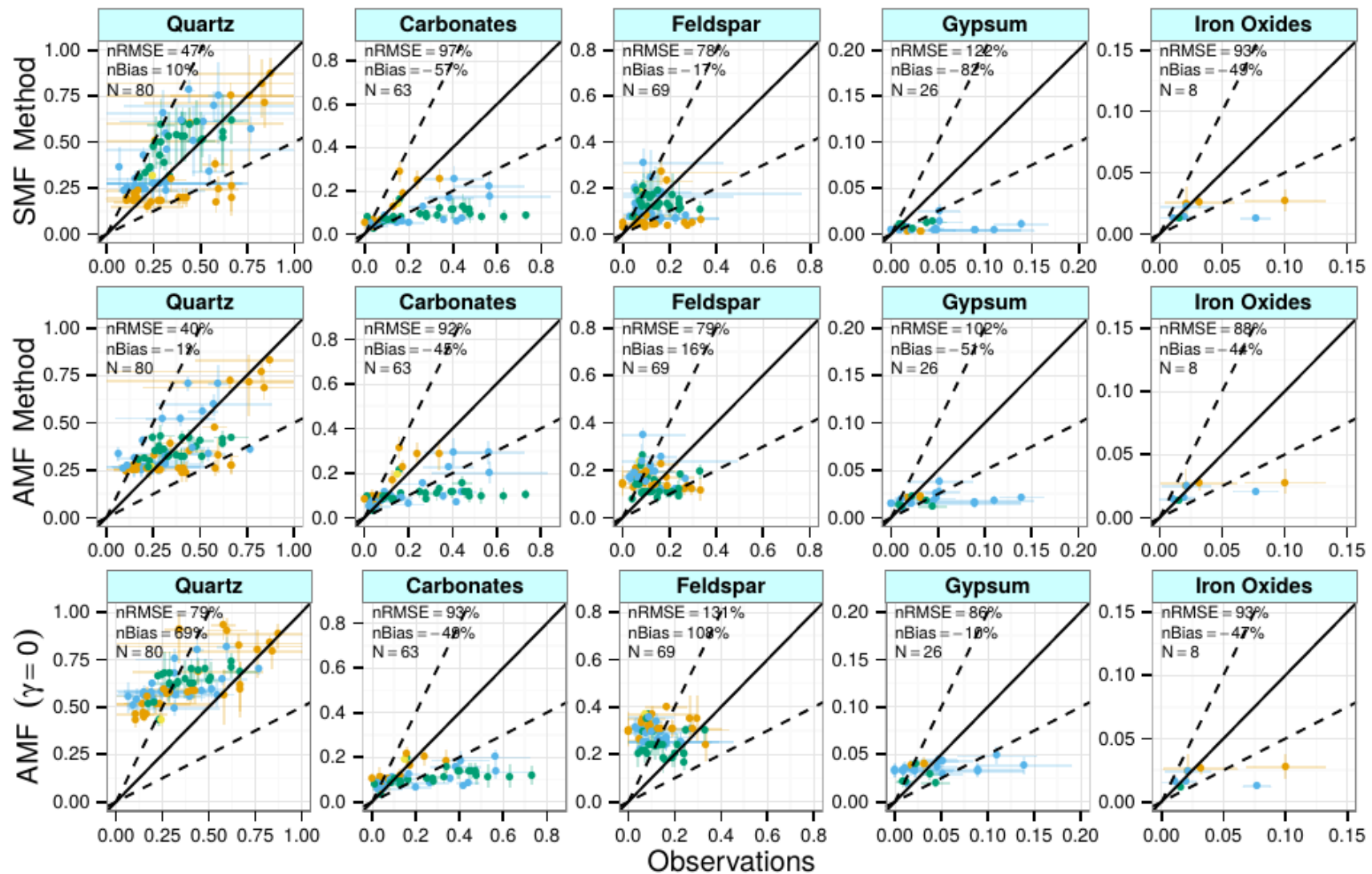
Evaluation - Bulk Dust: Phyllosilicates

Fractions of Minerals in Dust – Simulations versus Observations – Bulk Dust



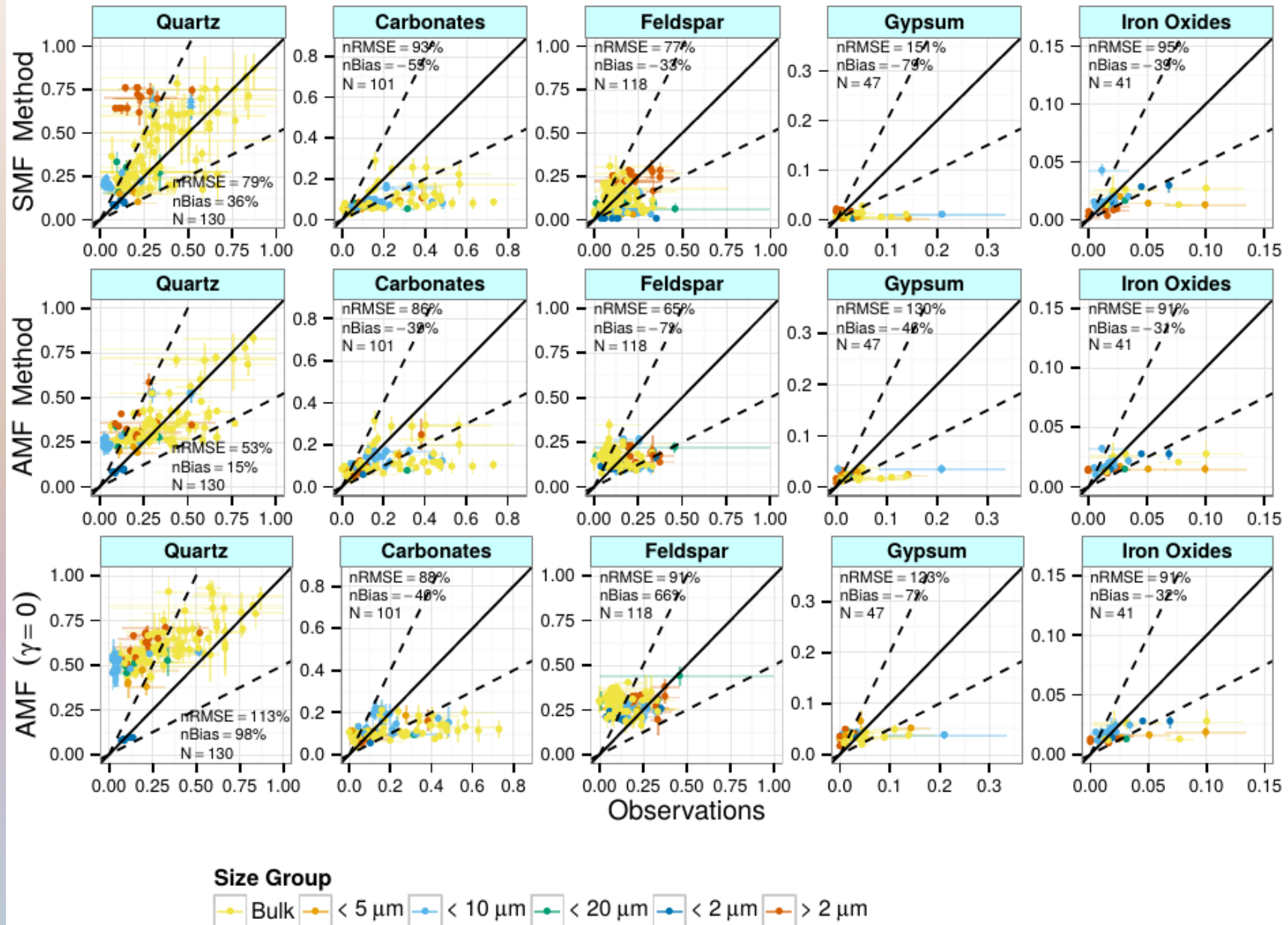
Evaluation - Bulk Dust: Other Minerals

Fractions of Minerals in Dust – Simulations versus Observations – Bulk Dust



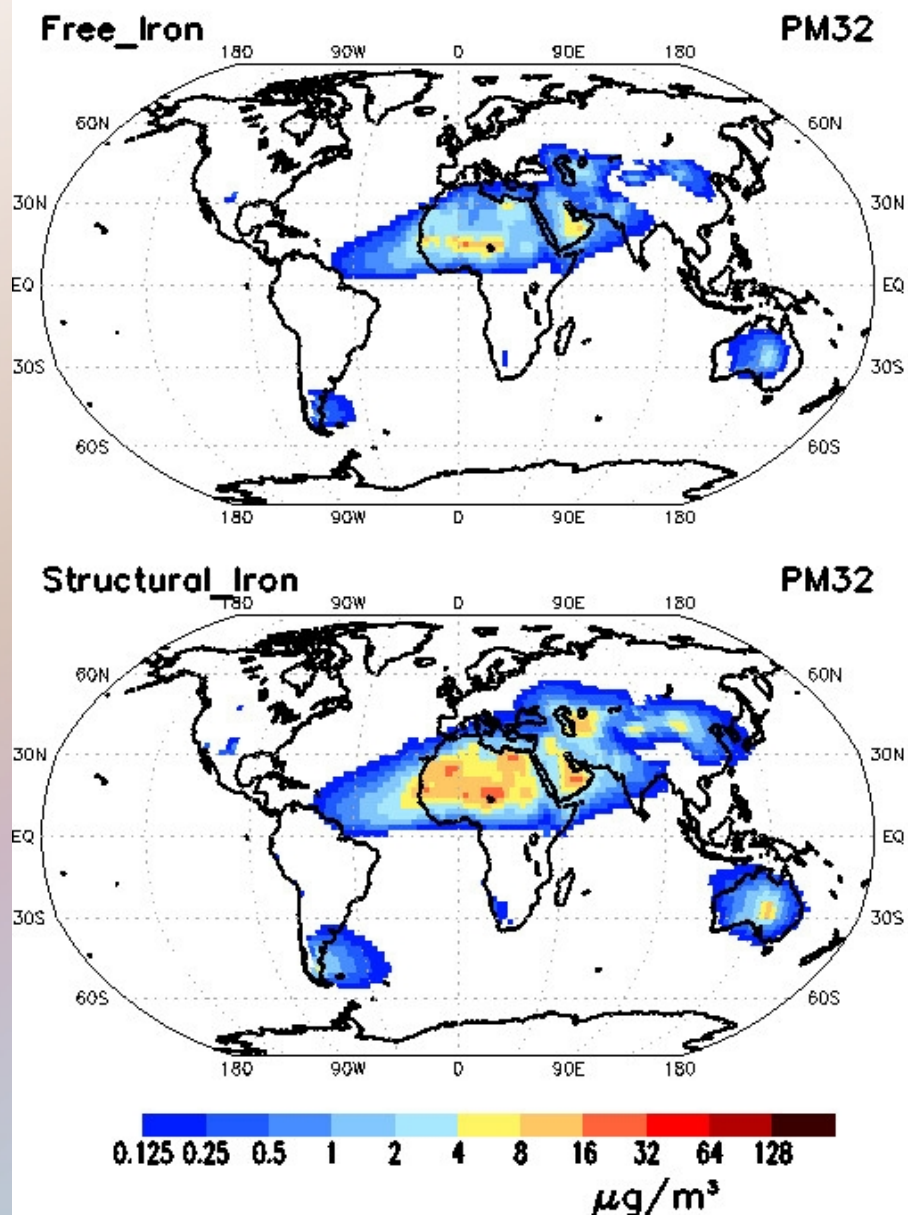
Evaluation – All Sizes Combined

Fractions of Minerals in Dust – Simulations versus Observations

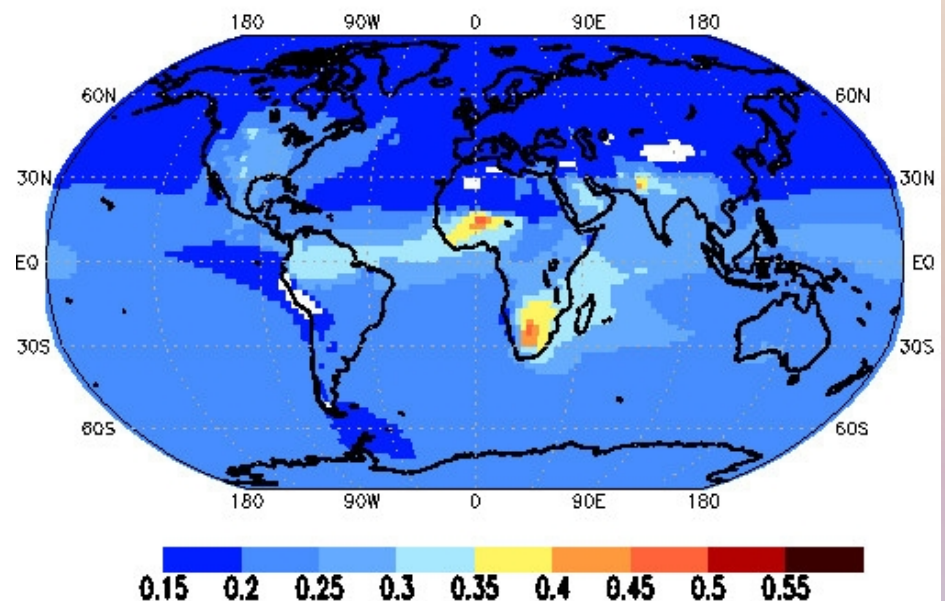


Example for Application: Free and Structural Iron

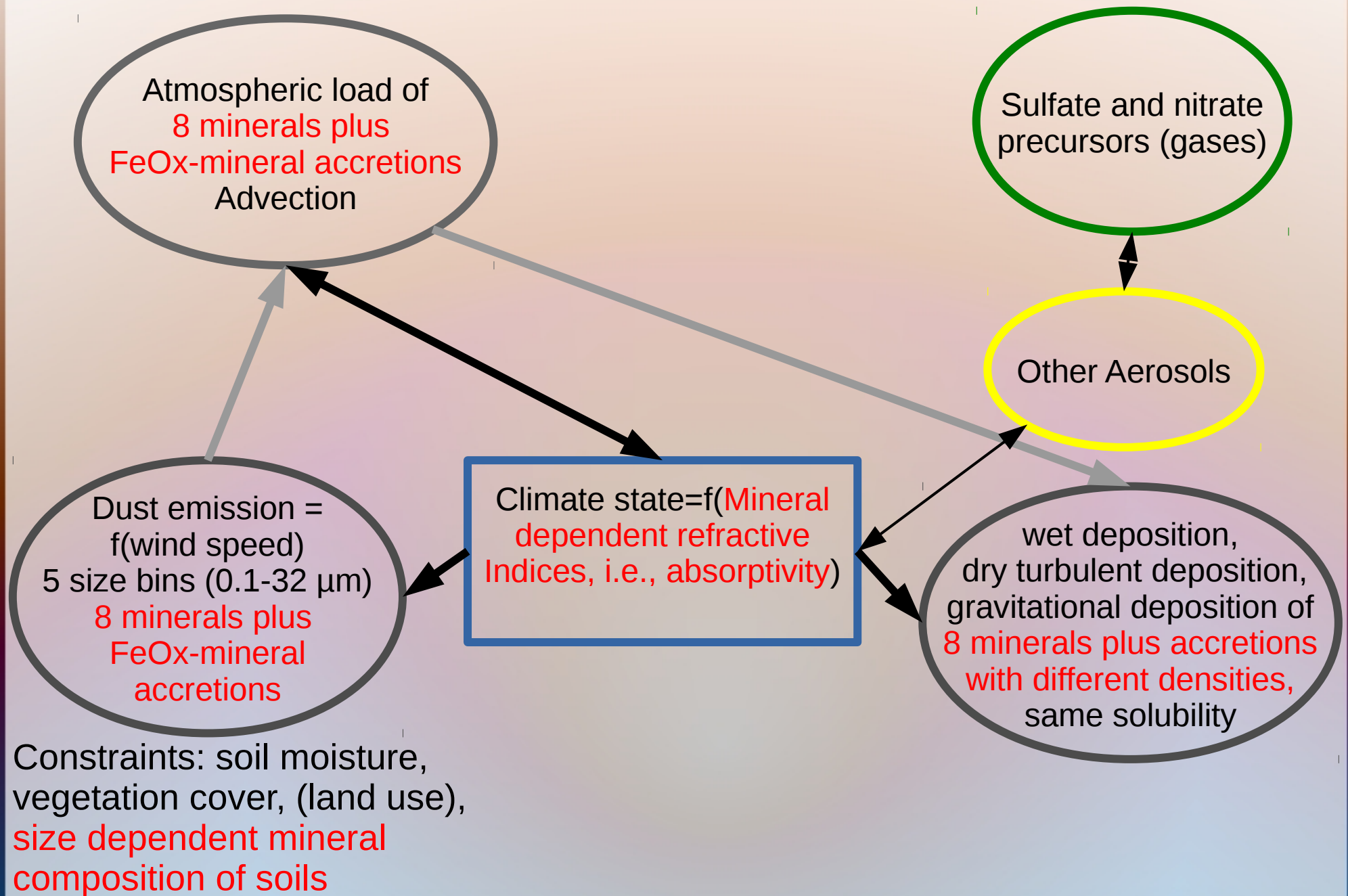
Surface Concentration of Free and Structural Iron
2002 to 2010 – Annual Mean



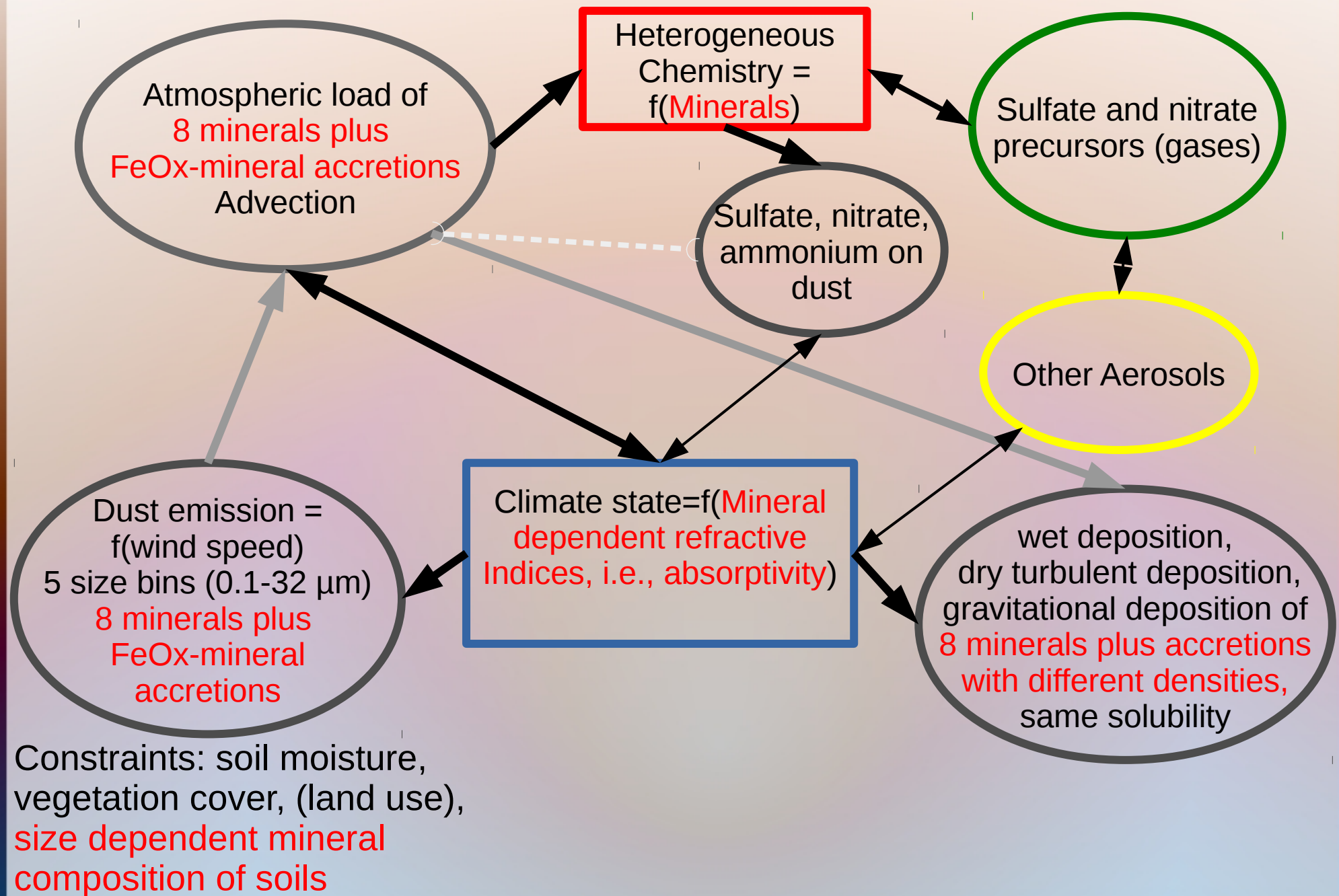
Free Iron Ratio from Surface Concentration
2002 to 2010 – Annual Mean



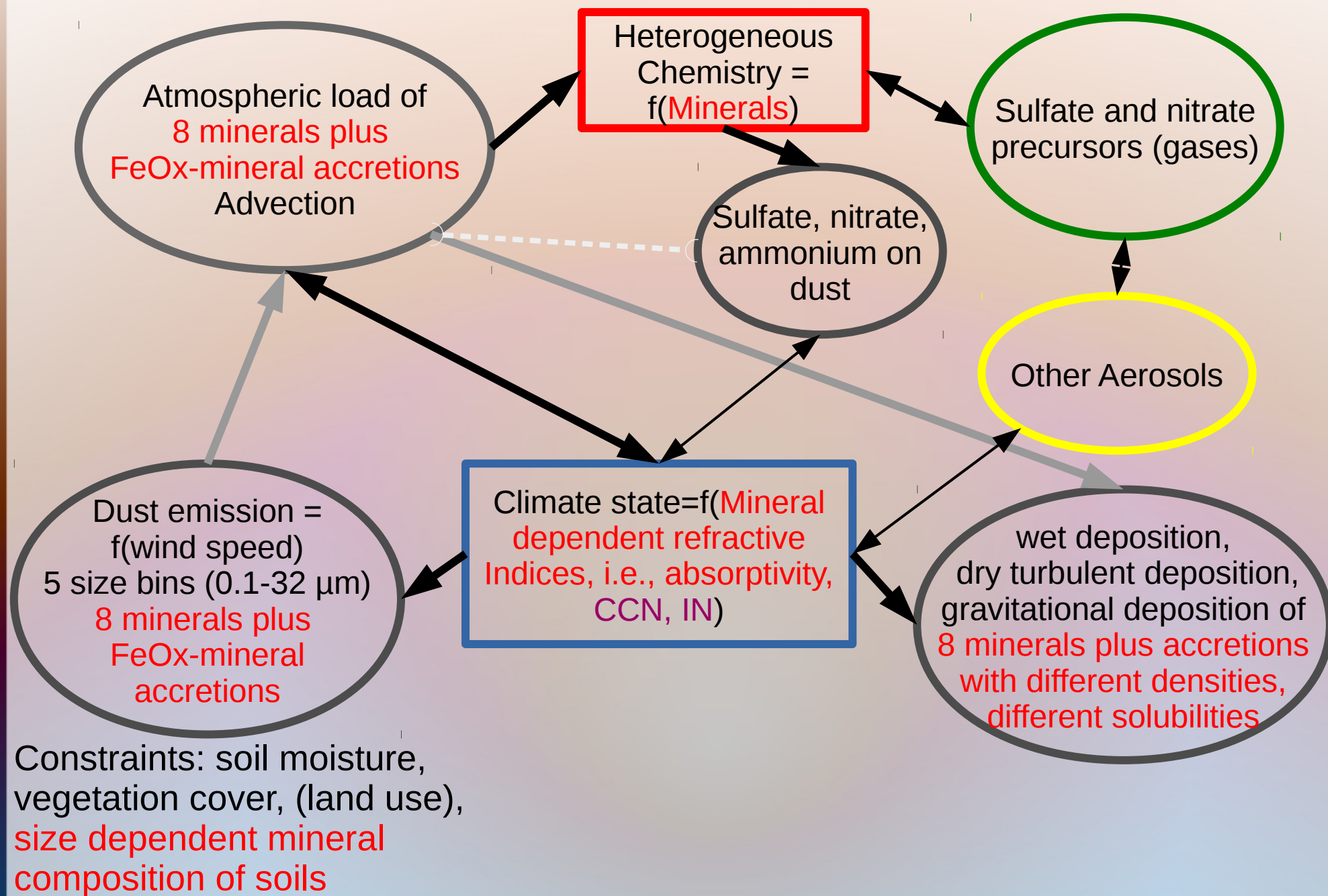
Next Steps: Radiative Effect of Minerals



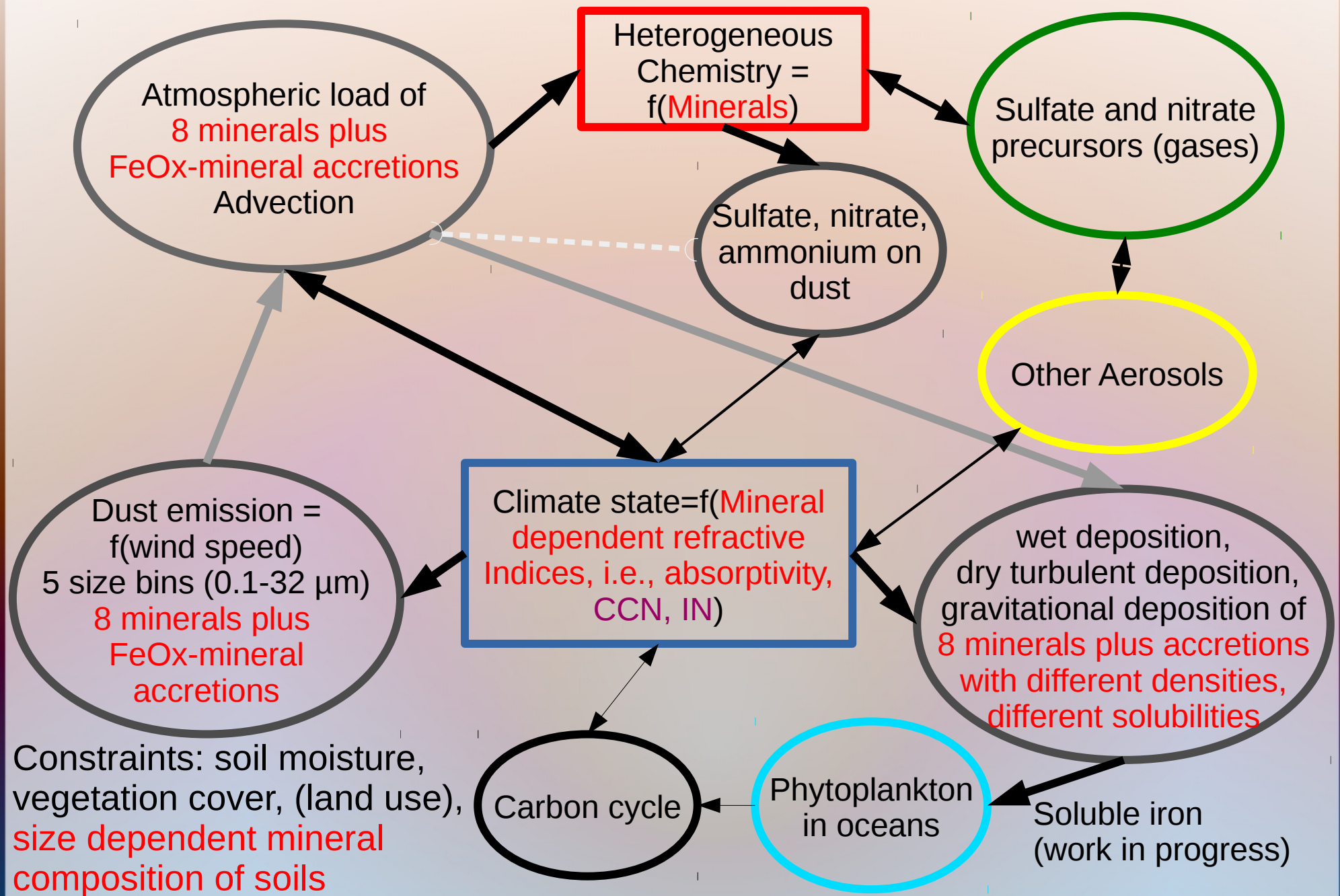
Next Steps: Heterogeneous Chemistry of Minerals



Near Future Work: Minerals as CCN and IN (collaborative)



Near Future Work: Iron Fertilization (collaborative)



Summary

- We have developed an improved approach for the derivation of the individual mineral species emission from soils, applying Claquin's Mean Mineralogical Table in combination with additional constraints from measurements by using aerosol size distributions for each mineral type
- We simulate emission, transport, and deposition of a mixture of non-accreted minerals and accretions of iron oxides with each of the other minerals in clay and silt size bins
- The agreement between data from measurements and model simulations is encouraging
- Claquin et al./Nickovic et al.'s MMT are very useful data. They have to be properly applied to derive the mineral fractions of dust aerosols

About The Future

In the future, for an improved validation we would need

- a) Data from routine measurements of the dust mineralogical composition over longer time periods
- b) The elemental composition of minerals, e.g., the amount of structural iron, to be able to use element data for the validation of the simulated mineral cycle
- c) More information about aggregation, especially of iron oxides in dust aerosols

Our To Do List (as collaborative efforts)

- a) Applying our methodology to the mineralogical table by Journet et al. ACP (2014)
- b) Sensitivity studies, variability of mineral composition
- c) Simulating iron fertilization (and, in turn, the effect on the carbon cycle), cloud condensation nuclei, ice nuclei, heterogeneous chemistry, radiative forcing of climate as processes dependent on the mineralogical composition of dust
- d) Potential for applications for paleo-climate studies

References:

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- Kandler, K., L. Schütz, C. Deutscher, M. Ebert, H. Hofmann, S. Jäckel, R. Jaenicke, P. Knippertz, K. Lieke, A. Massling, A. Petzold, A. Schladitz, B. Weinzierl, A. Wiedensohler, S. Zorn, and S. Weinbruch (2009), Size distribution, mass concentration, chemical and mineralogical composition and derived optical parameters of the boundary layer aerosol at Tinfou, Morocco, during SAMUM2006, *Tellus*, 61B(1), 32–50, doi:10.1111/j.1600-0889.2008.00385.x.
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- Nickovic, S., A. Vukovic, M. Vujadinovic, V. Diurdevic, and G. Pejanovic (2012), *Atmos. Chem. Phys.*, 12(2), 845-855, doi: 10.5194/acp-12-845-2012.
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